

FARCOS

Femtoscope Array for Correlations & Spectroscopy

- Overview of project and physics inputs
- Present status (mechanics, electronics) and perspectives for the use with GET electronics

G. Cardella, INFN-Catania

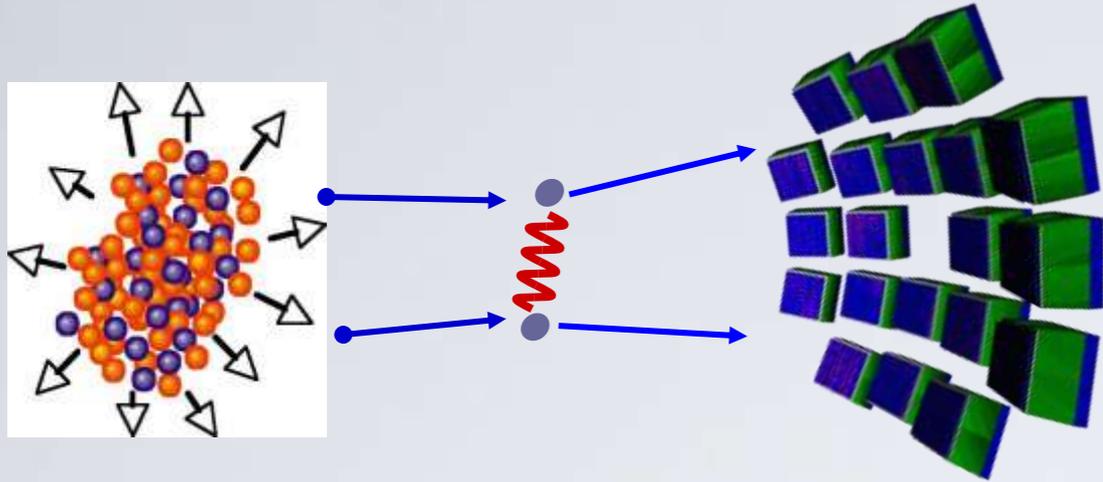
EXOCHIM collaboration, INFN (LNS-Ct-Me-Mi-Na)

GANIL, GEM-UHU, open collaboration

Physics topics

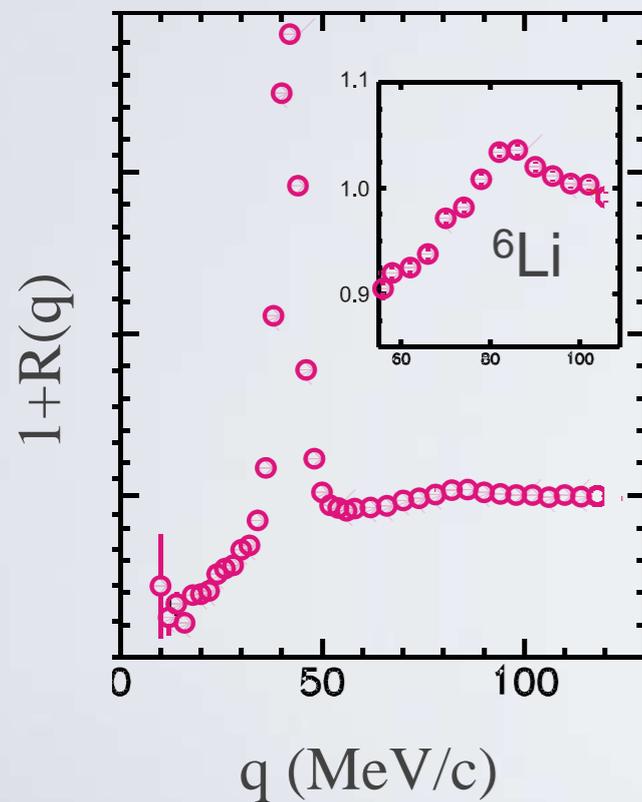
1. Heavy-ion collisions (stable and RI beams)
 - Dynamics (HBT, Femtoscopy)
 - Low & Intermediate energies: fusion, fission, DIC, Symmetry energy, Emission time-scales, probes of reaction models
 - Multi-particle correlation spectroscopy (MPCS)
 - cluster states
2. Direct reactions with RIBs
 - Inverse and direct kinematics
 - Nuclei close to drip lines

Relevance of angular resolution



$$1 + R(q) = k \times \frac{SY_{coinc}(\vec{p}_1, \vec{p}_2)}{SY_{evt.mixing}(\vec{p}_1, \vec{p}_2)}$$

deuteron-alpha

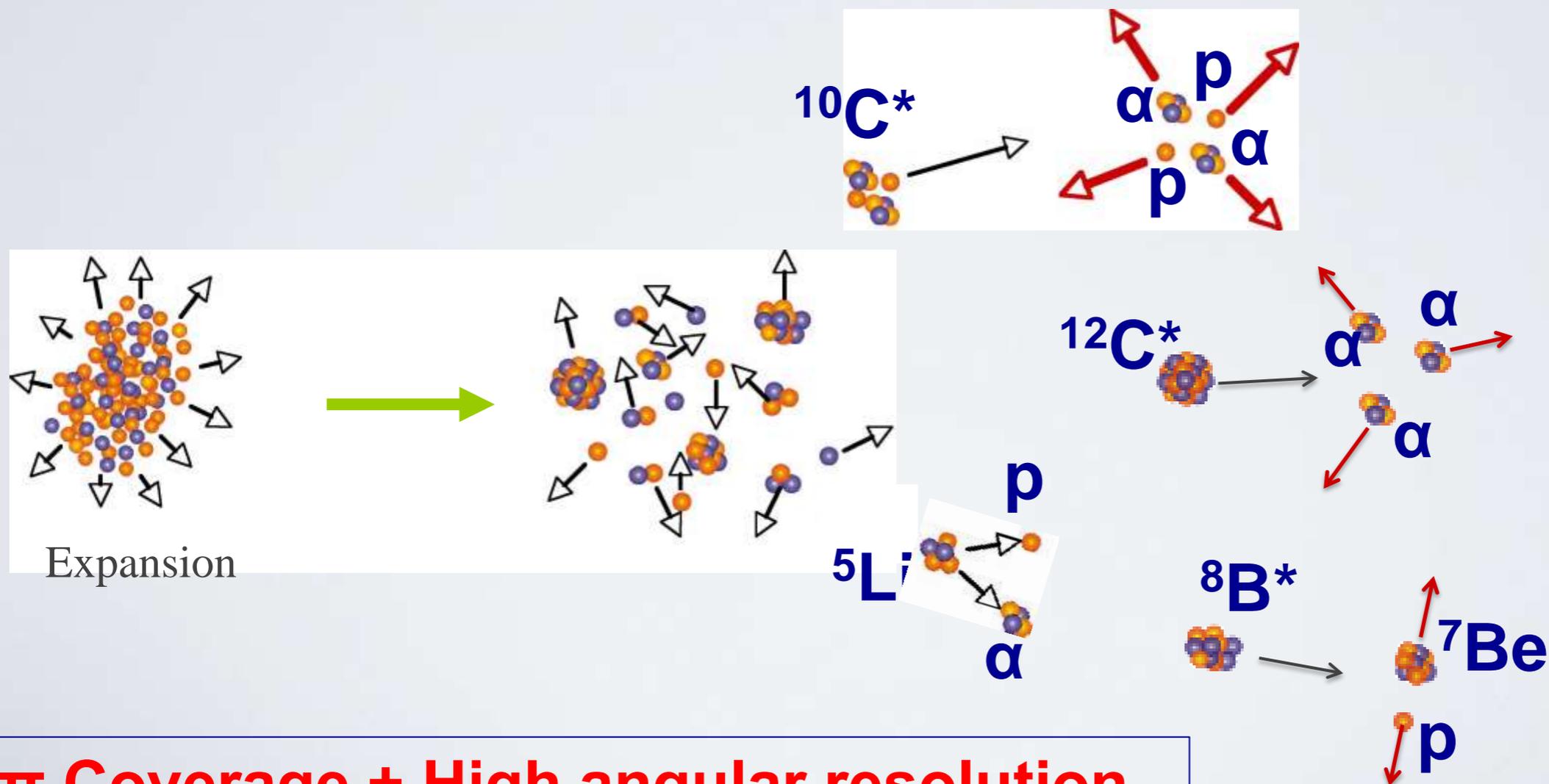


$$q = m|\vec{v}_1 - \vec{v}_2|$$

High angular resolution required:
 → Measurements at low q and sharp resonances

Multi-particle decay spectroscopy

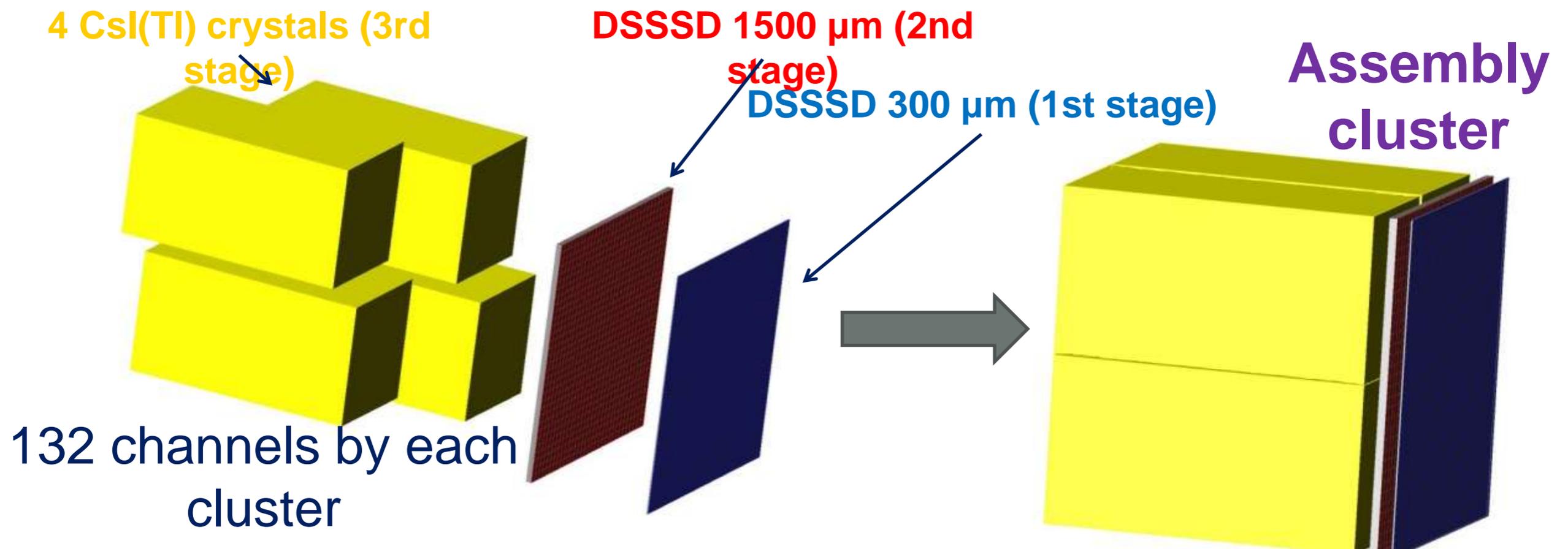
Using heavy-ion collisions as an explorative spectroscopy tool:
several unbound species in one single experiment



4 π Coverage + High angular resolution

FARCOS TELESCOPE – phase 1

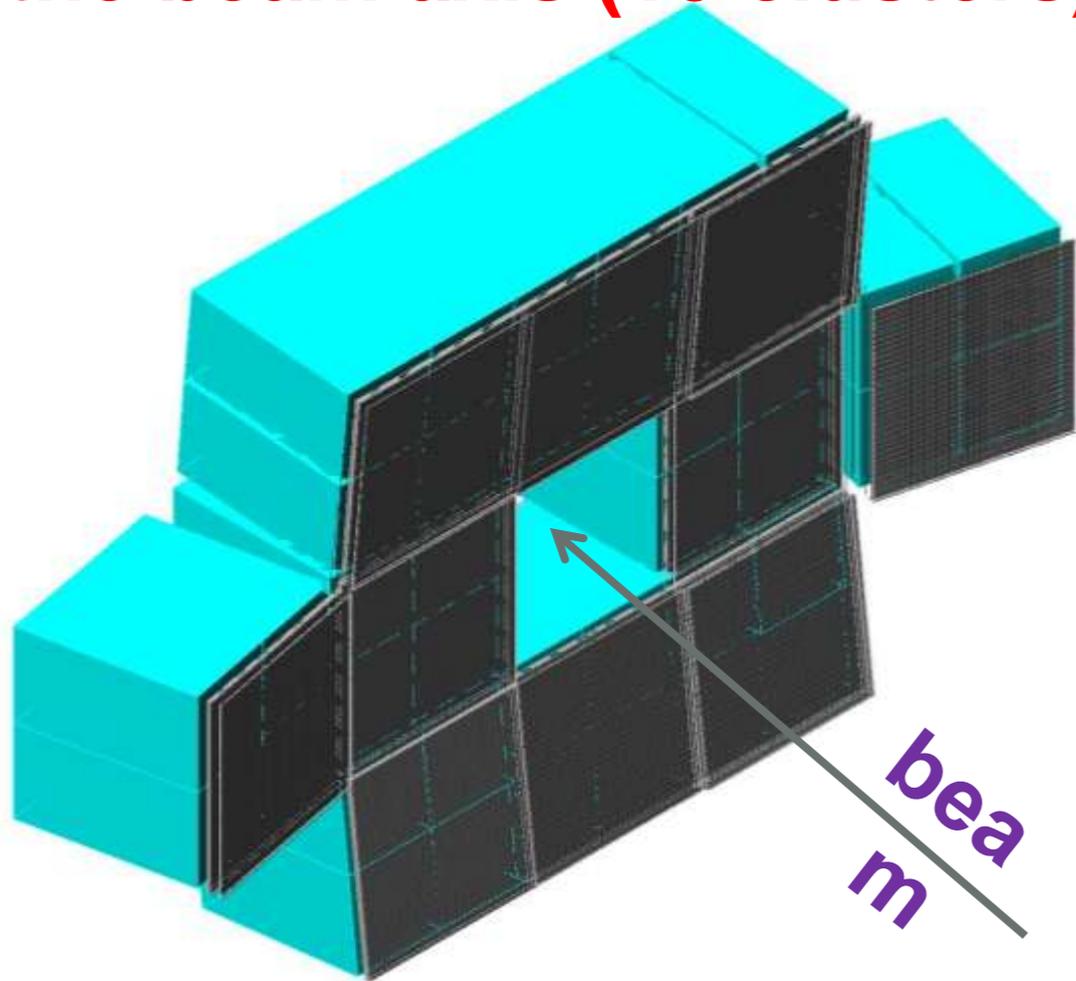
- Based on $(62 \times 64 \times 64 \text{ mm}^3)$ clusters
- **1** square $(0.3 \times 62 \times 62 \text{ mm}^3)$ DSSSD 32+32 strips
- **1** square $(1.5 \times 62 \times 62 \text{ mm}^3)$ DSSSD 32+32 strips
- **4** $60 \times 32 \times 32 \text{ mm}^3$ CsI(Tl) crystals



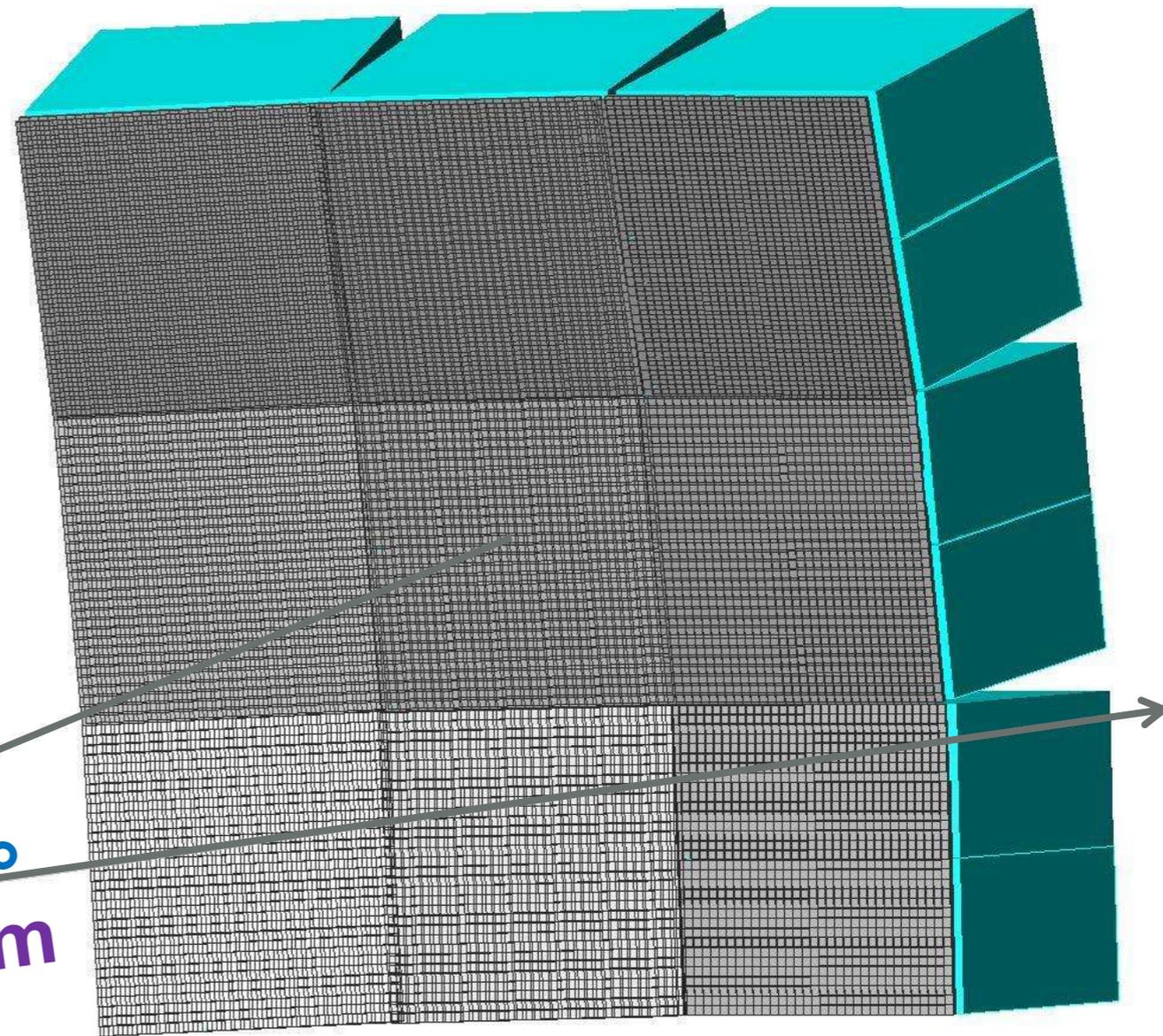
Fully reconfigurable (more Si layers, neutron and γ -detection, ..)

Geometric flexibility

Cross geometry centered on the beam axis (10 clusters)

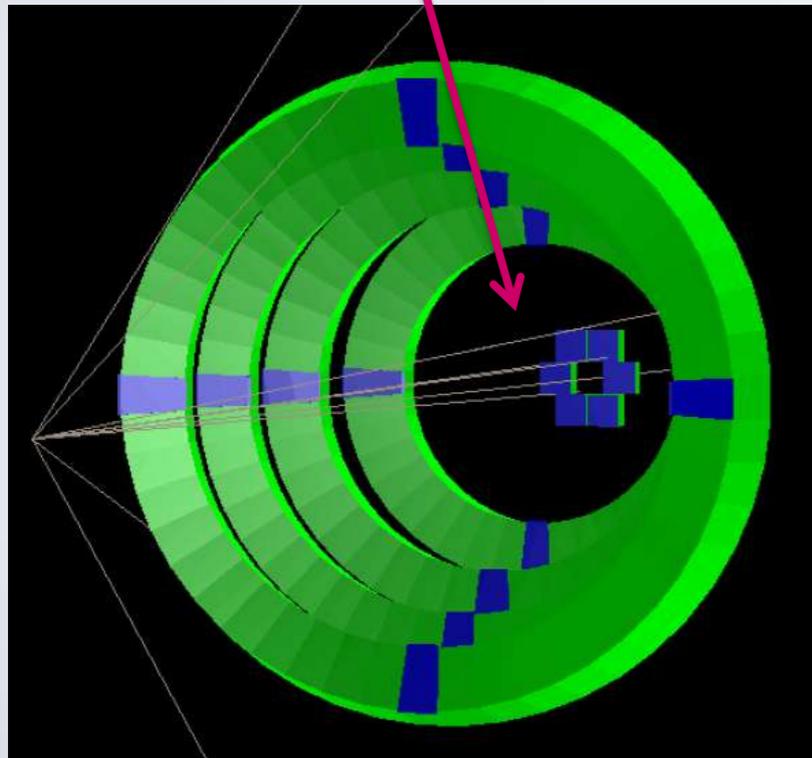
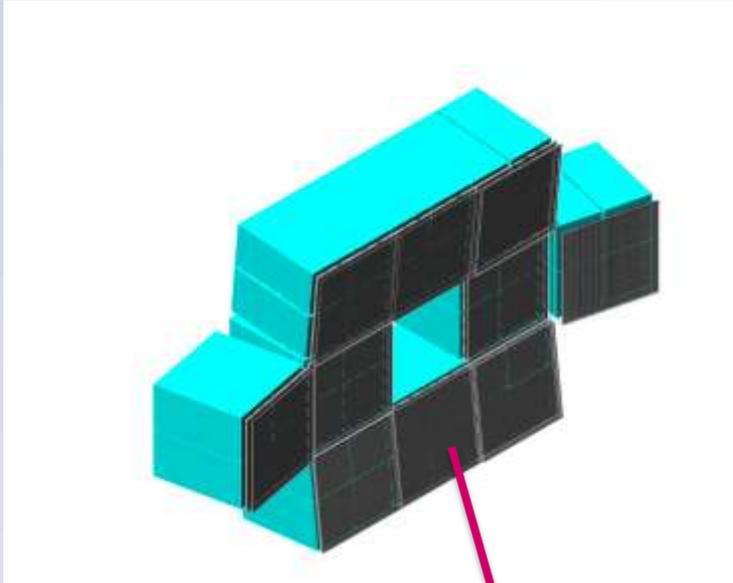


Wall geometry placed at 45° from beam axis (9 clusters)



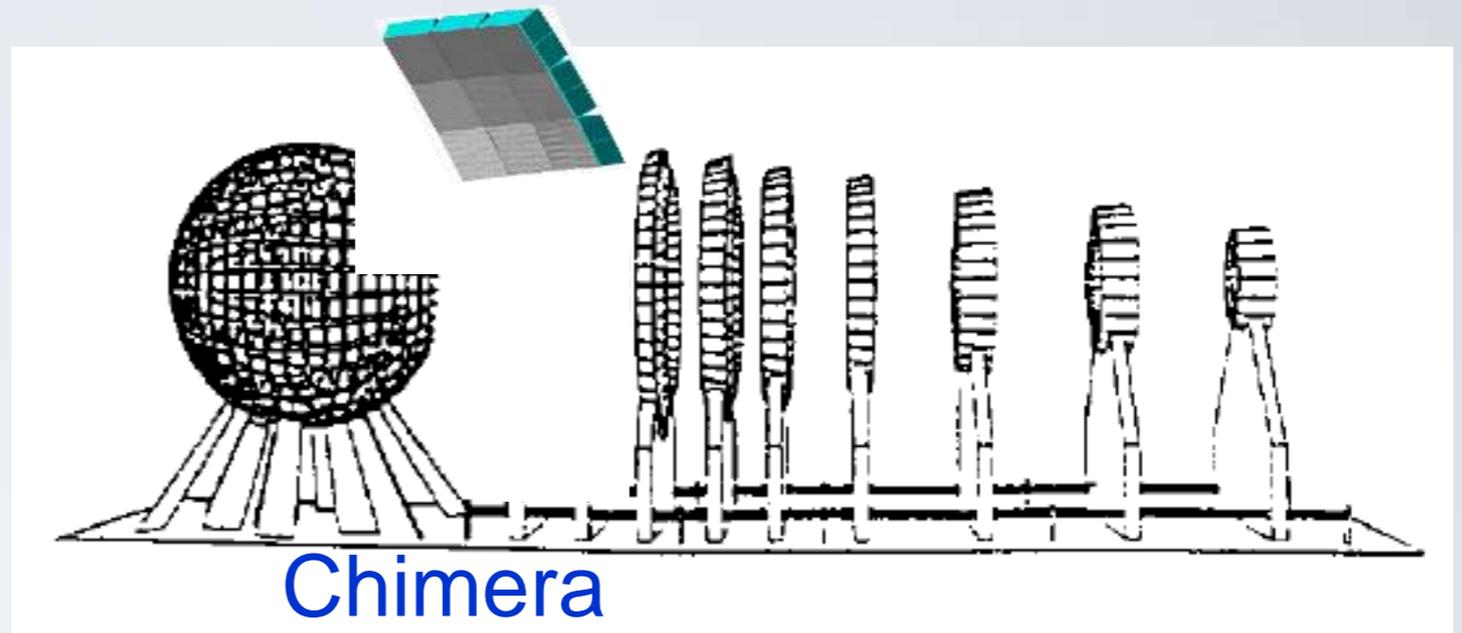
Transportability
Modularity

Planned geometries



Chimera rings

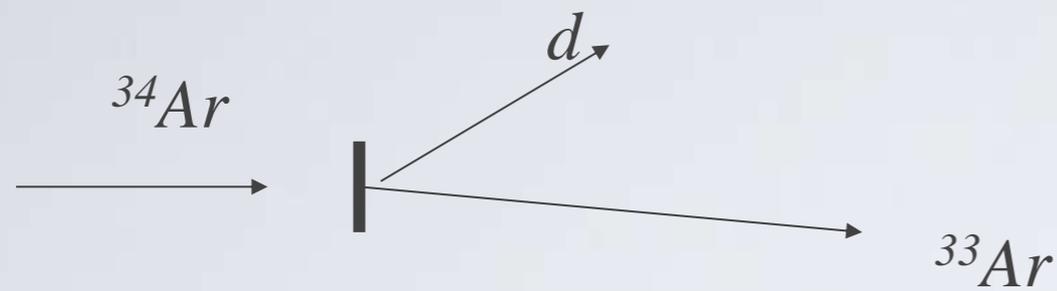
Farcos



- Correlations in central collisions (cluster states, HBT, Asy-EoS)
- Correlations in Quasi-Projectile breakup
- Spectroscopy with radioactive beams

Direct reactions with exotic beams

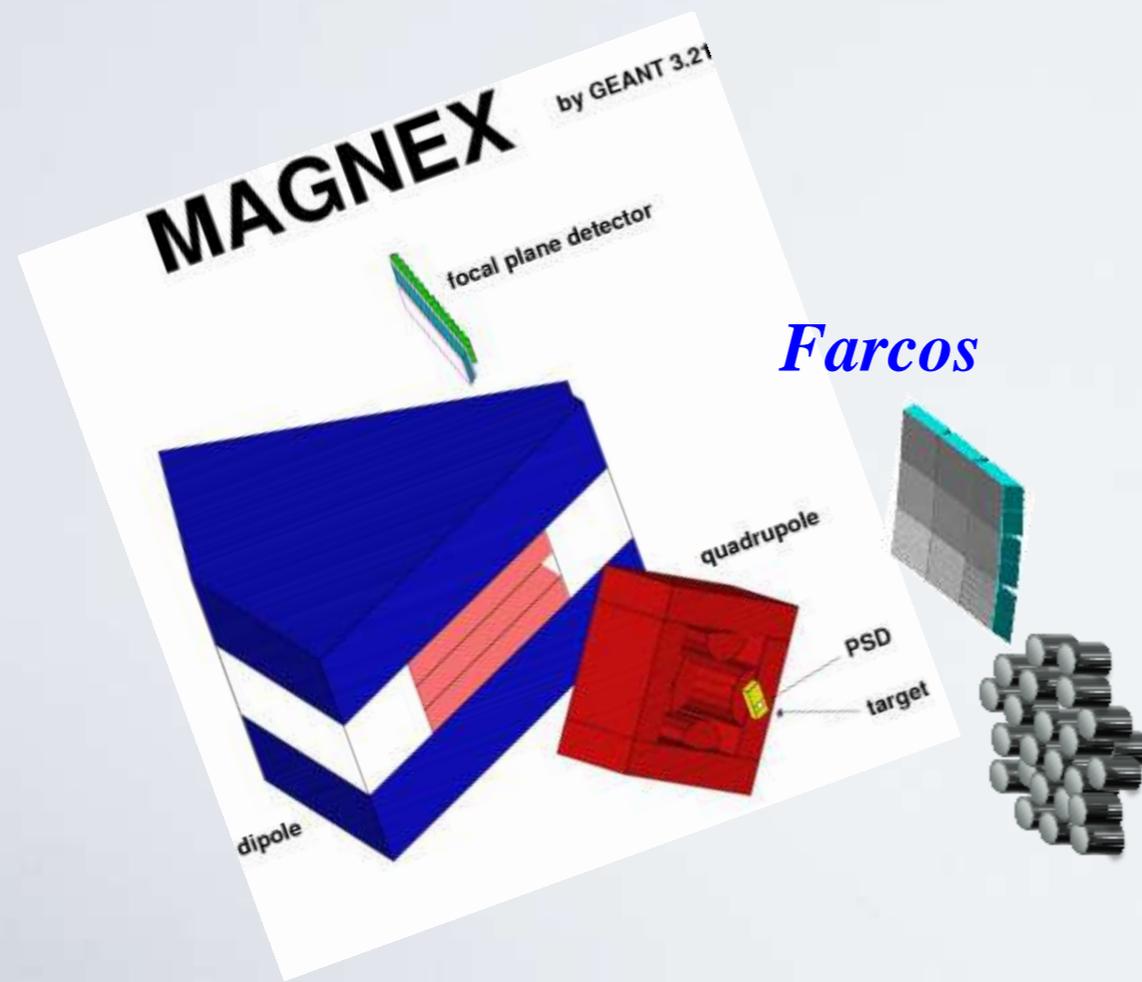
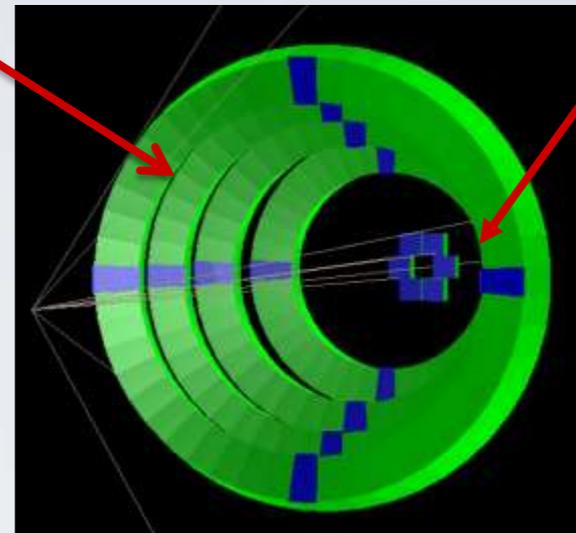
FRIBS exotic beams



Chimera

(d)

Farcos (^{33}Ar residue)



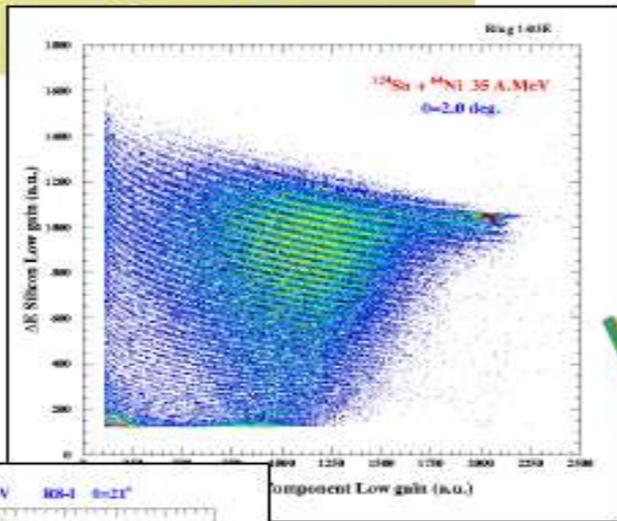
Coupling to spectrometers

Neutron detectors

Required identification performances (Chimera-like)

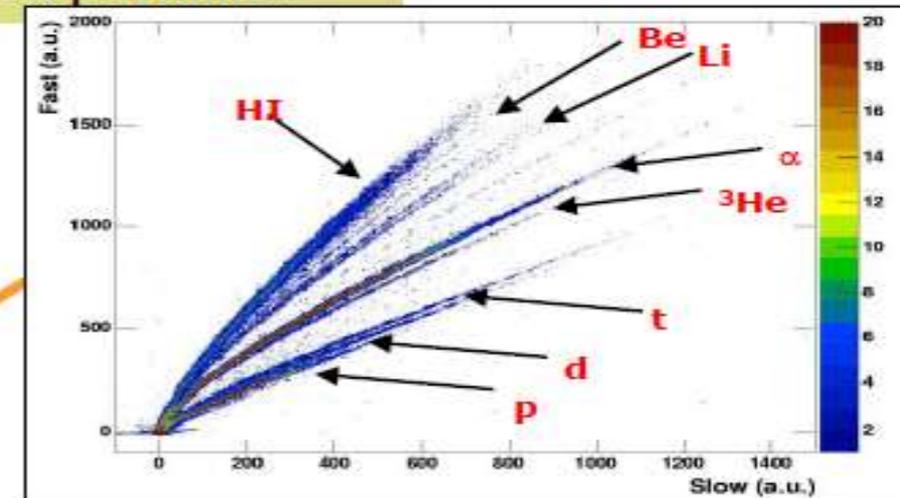
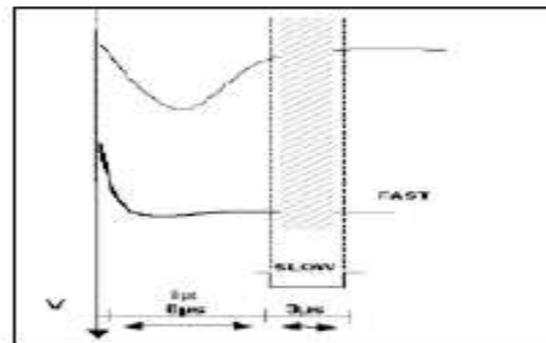
$\Delta E(\text{Si})-E(\text{CsI})$

Charge Z for particles punching through the Si detector



PSD in CsI(Tl)

Z and A for light charged particles

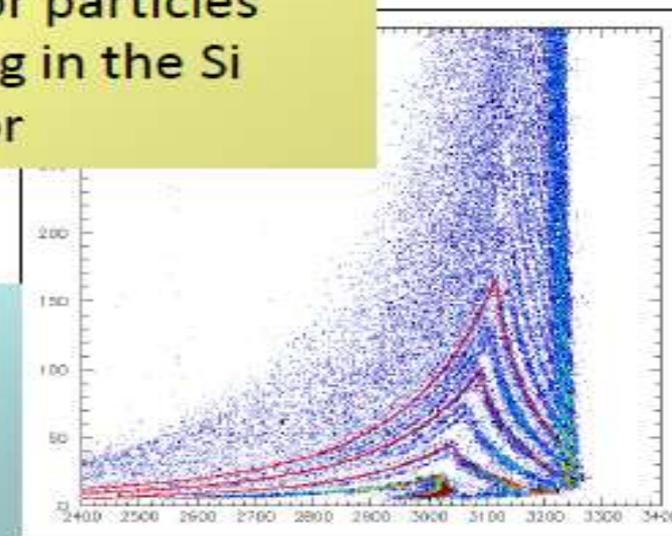


$\Delta E(\text{Si})-\text{ToF}$

Mass for particles stopping in the Si detector

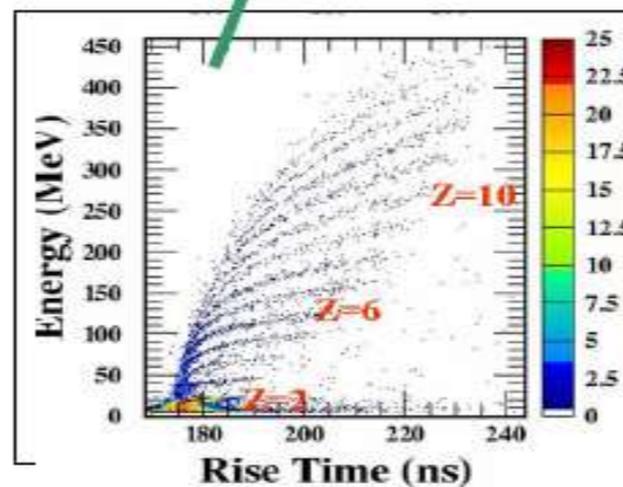
Si
~300 μm

CsI(Tl)
3-12 cm



$\Delta E(\text{Si})-E(\text{CsI})$

Charge Z and A for light ions ($Z < 9$) punching through the Si detector

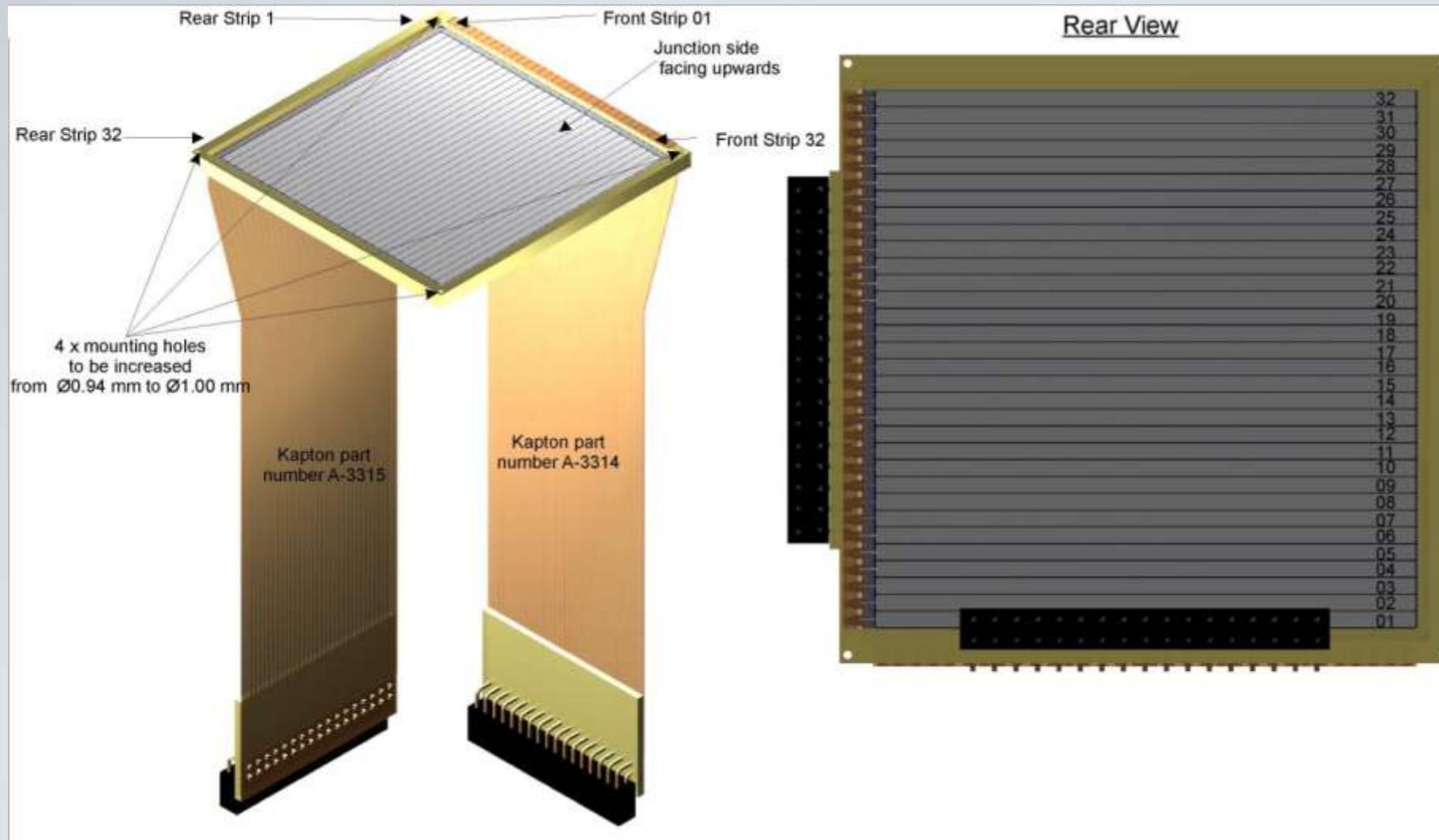


$E(\text{Si})-\text{Rise time}$

Charge Z for particle stopping in Si detectors

Large dynamical range

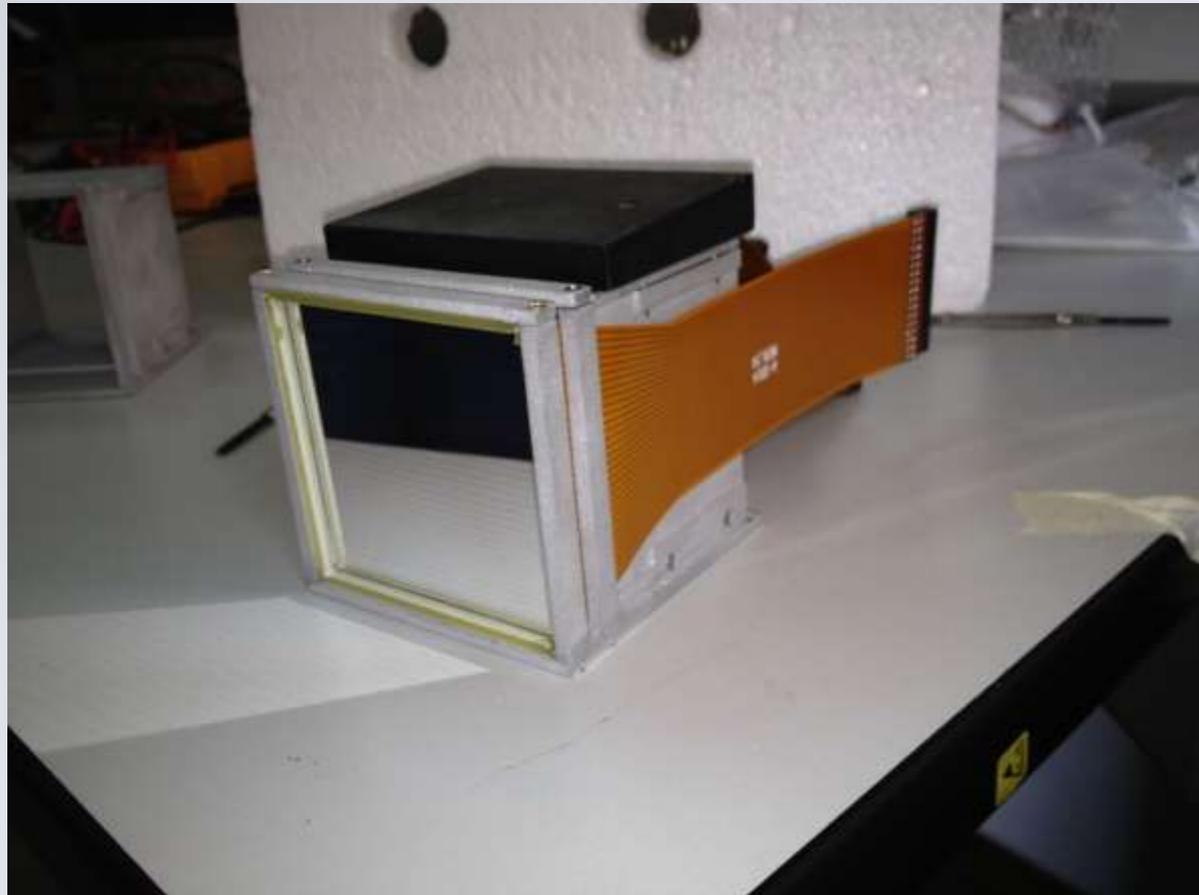
FARCOS detectors



- Double-Sided Silicon Strip Detectors
- 300 μm and 1500 μm
- kapton cable and 2x32pin connectors

- Highly homogenous CsI(Tl) crystals
 - Wrapping: 0.12mm thick white reflector +50 μm aluminized Mylar.
- 2 μm thick aluminized Mylar window at the entrance (0.29 g/cm^2)
- Read-out by photo-diodes (300 μm)

First prototype modules built

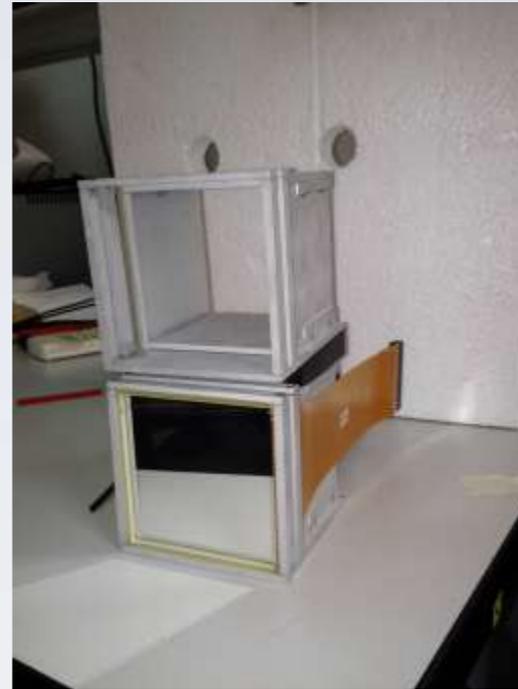


4 clusters ready
Experiments in Spring 2013 at
LNS approved



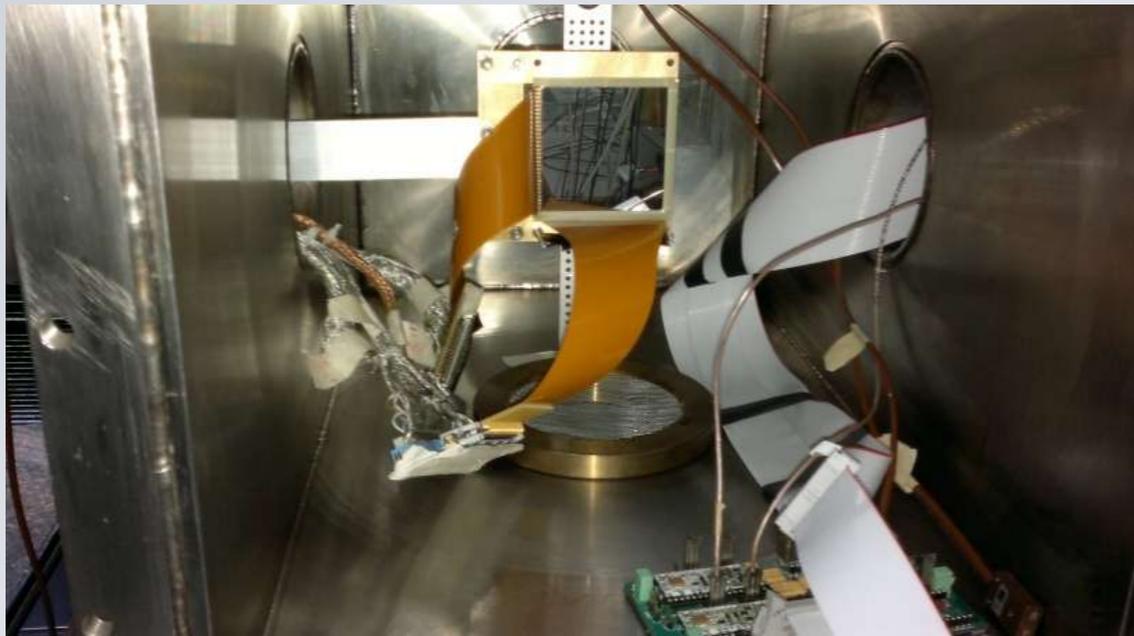
Compact preamplifiers – Phase 1

32 channels Hybrid charge preamplifiers in a volume of about 8cm x 10cm x 2mm



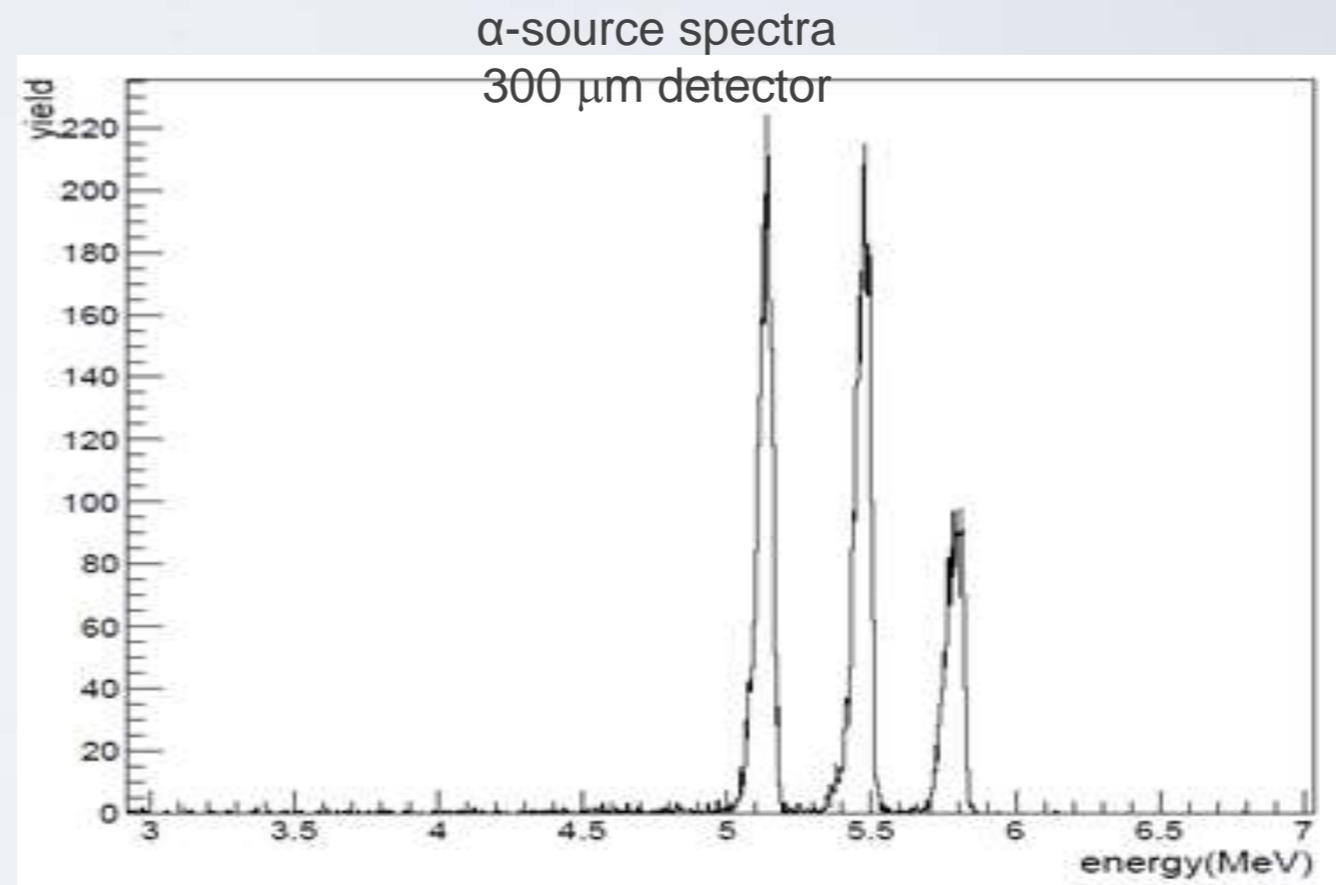
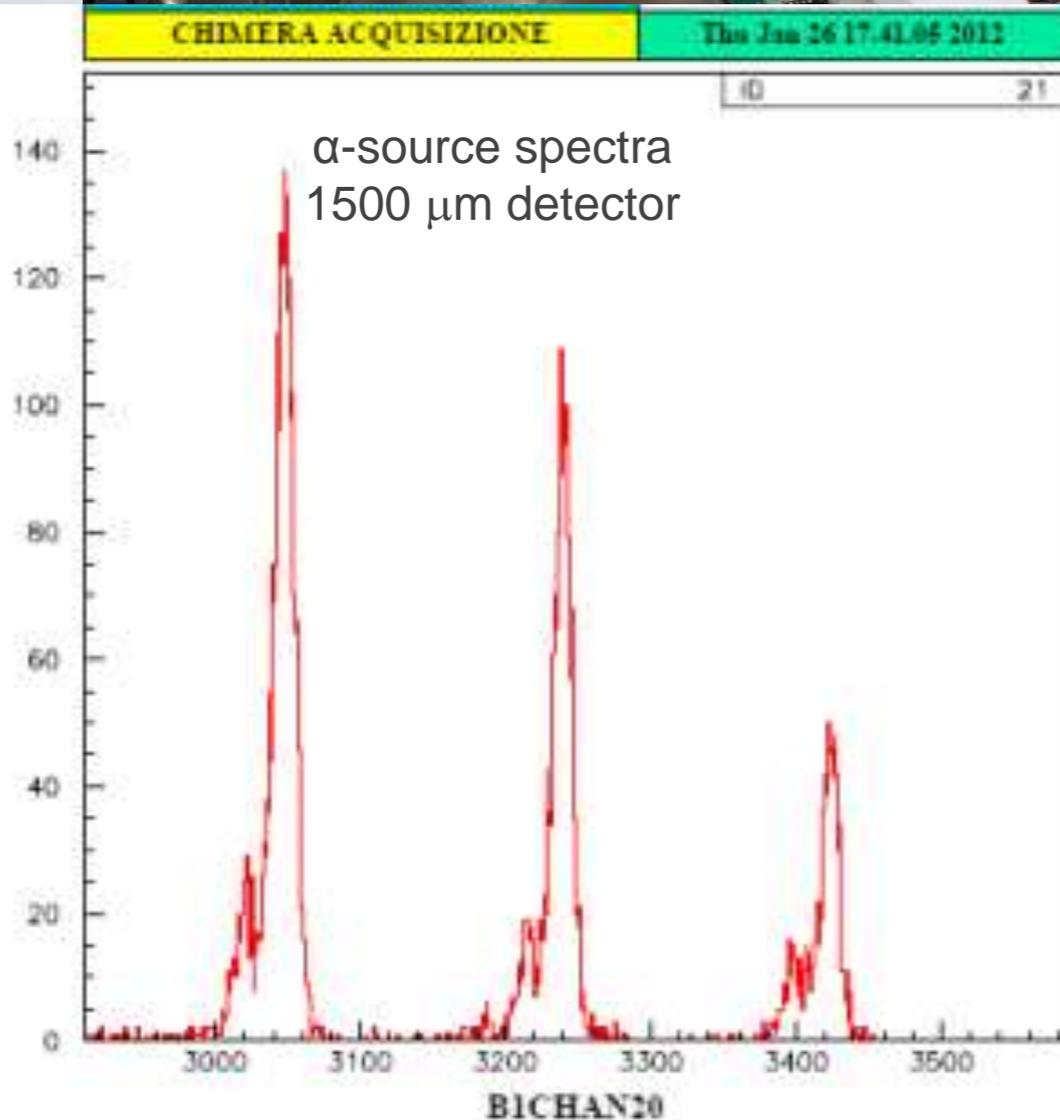
- **Low power consumption:** ~750 mW pwe 32 channels (simplify cooling operations)
- **Rise-Time (pulser):** ~ 3-7 nsec for $C_{input}=0-100pF$
- **Energy resolution (pulser)** ~ 4-5 KeV for $C_{input}=0-100pF$
- **Available with several sensitivities** (5, 10, 45, 100 mV/MeV...)

Si-strips α -source tests



INFN-preamplifiers and Mesytec provide comparable E resolution (about 1% for the 300 μm strip .5% for 1500 μm)

INFN more compact, faster, less power consumption

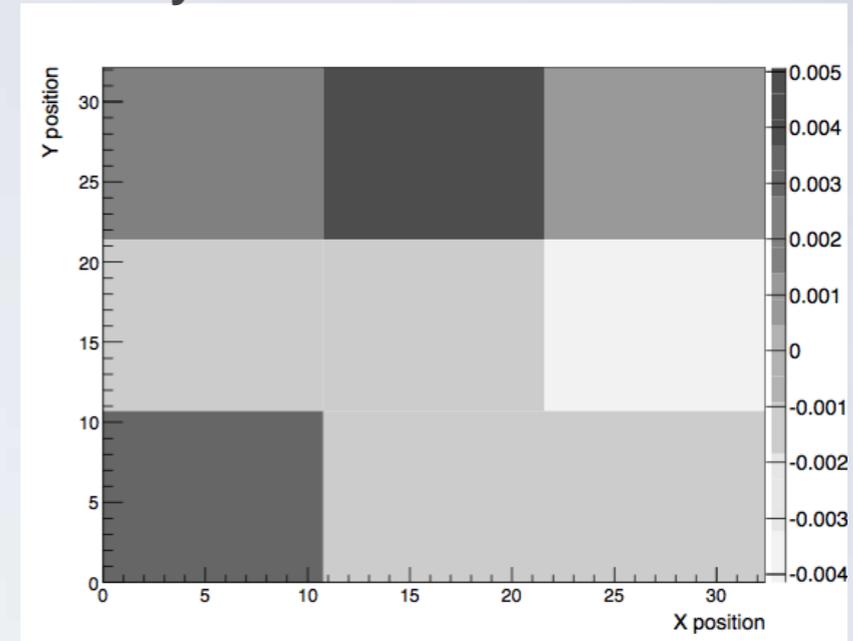


CsI(Tl) uniformity tests with α -source

Source moved along silicon face (front and back)



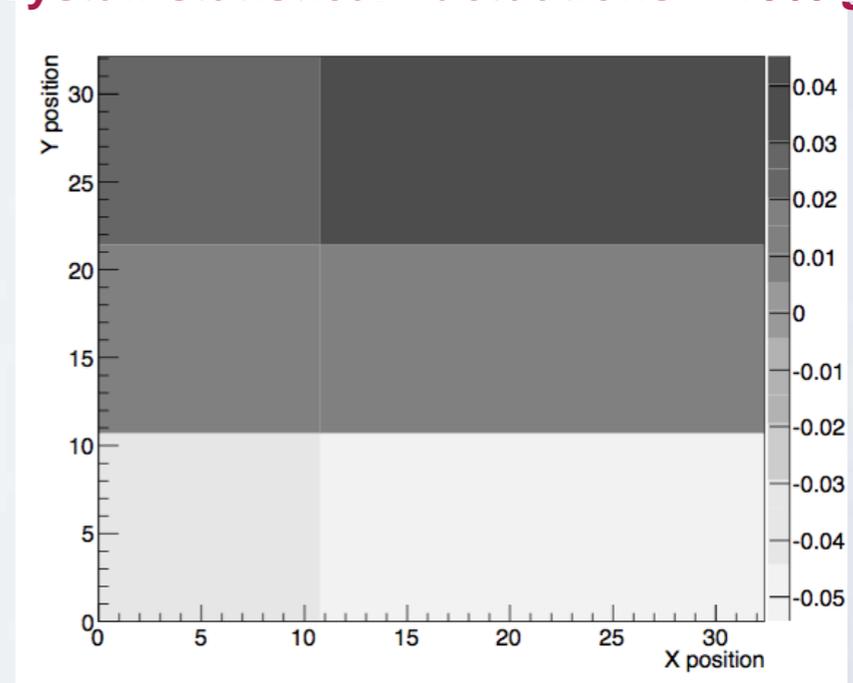
“Good” crystal: statistical fluctuations $< 1\%$



Non-uniformities (if systematic) can be corrected thanks to position sensitivity of silicon strips (for higher energy particles)

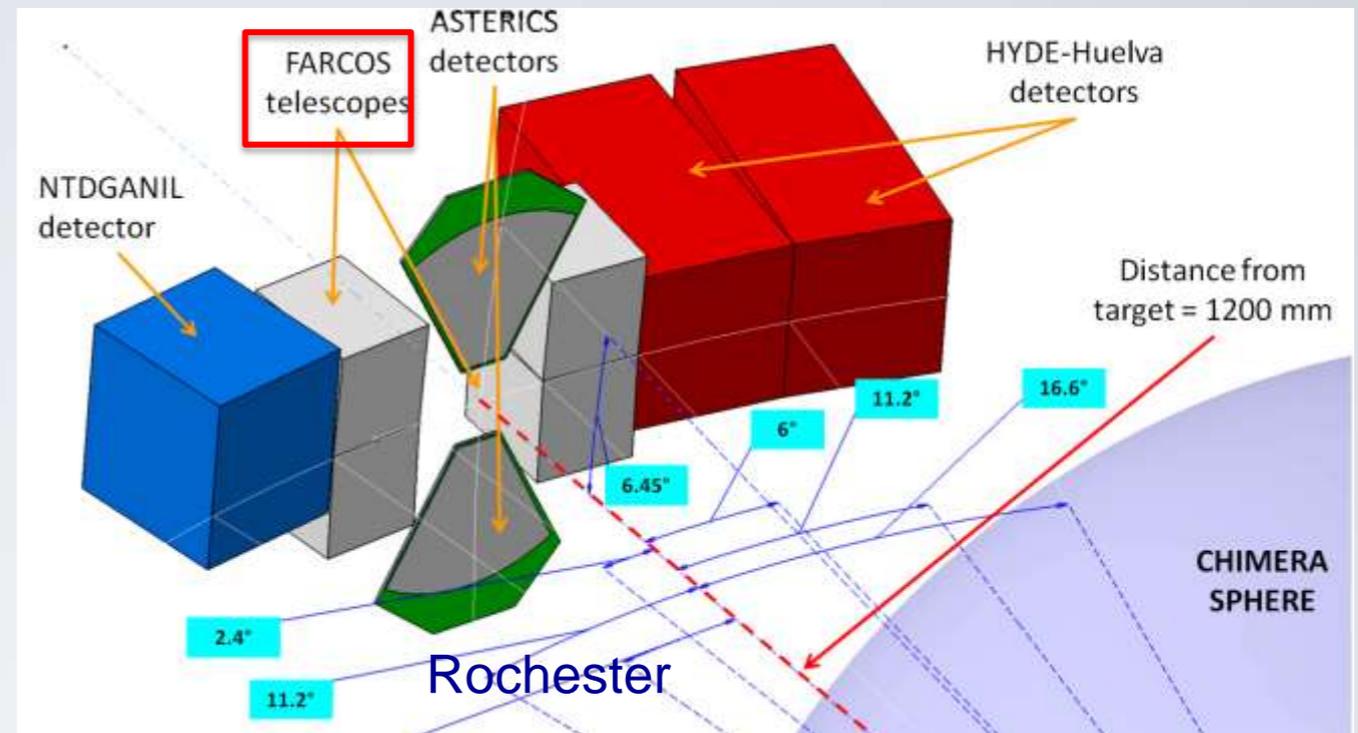
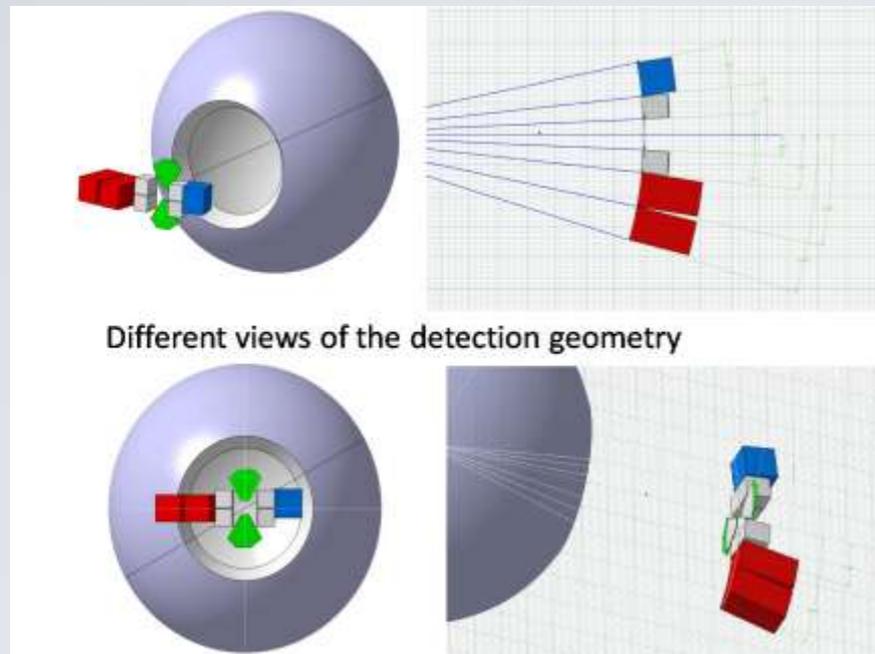
Improve resolution in beam experiments

“Bad” crystal: statistical fluctuations $< 10\%$ gradient



First test beams – July 2012 @ LNS

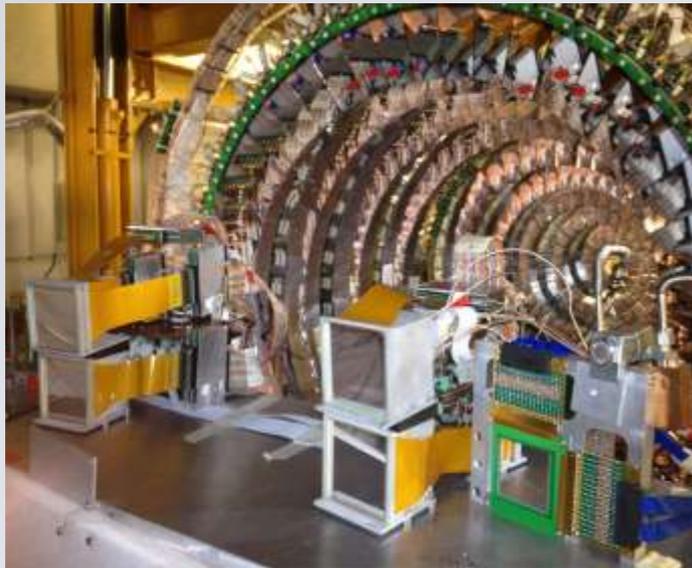
Chimera sphere view



p, α , Ne beams $E/A=40, 20$ MeV - Elastic scattering and Transfer reactions on various targets

- **CsI(Tl) uniformity**
- **Silicon resolution**
- **Integrating DAQ into Chimera system**

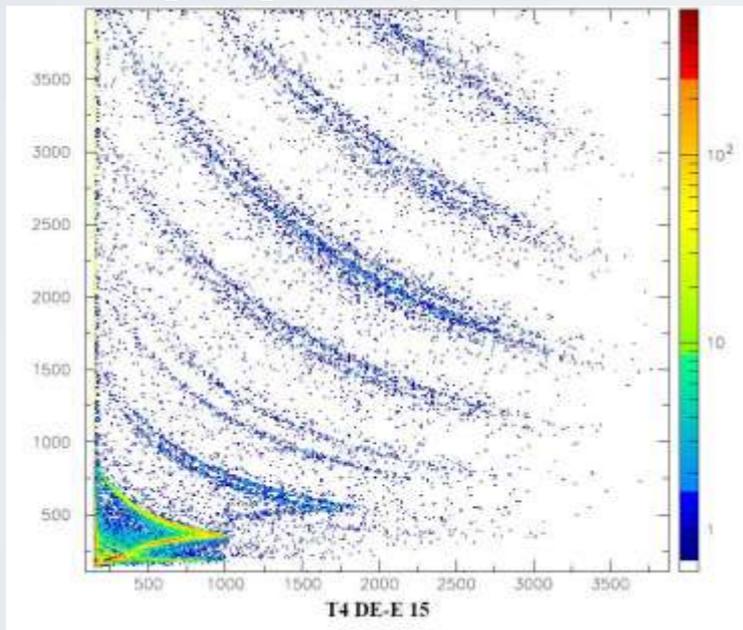
Mounting of prototype telescopes



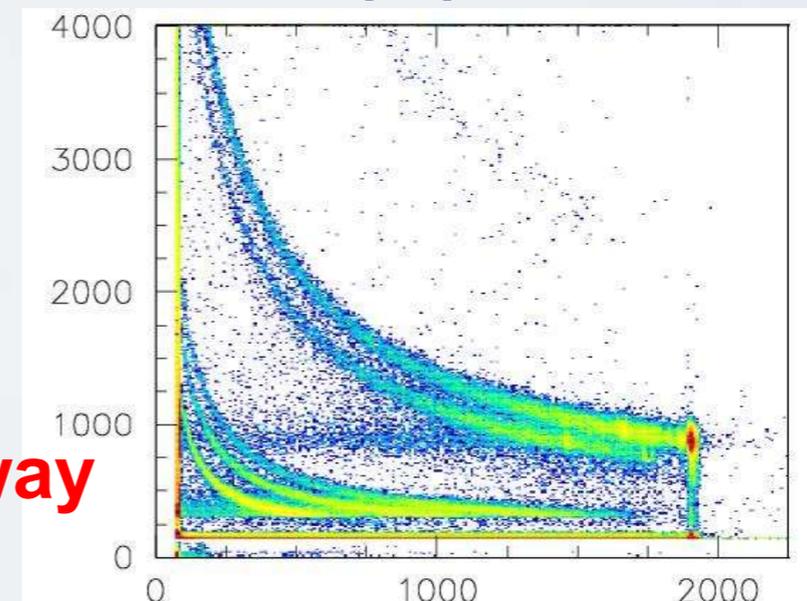
Pre-amplifier
boxes



Si-Si matrix



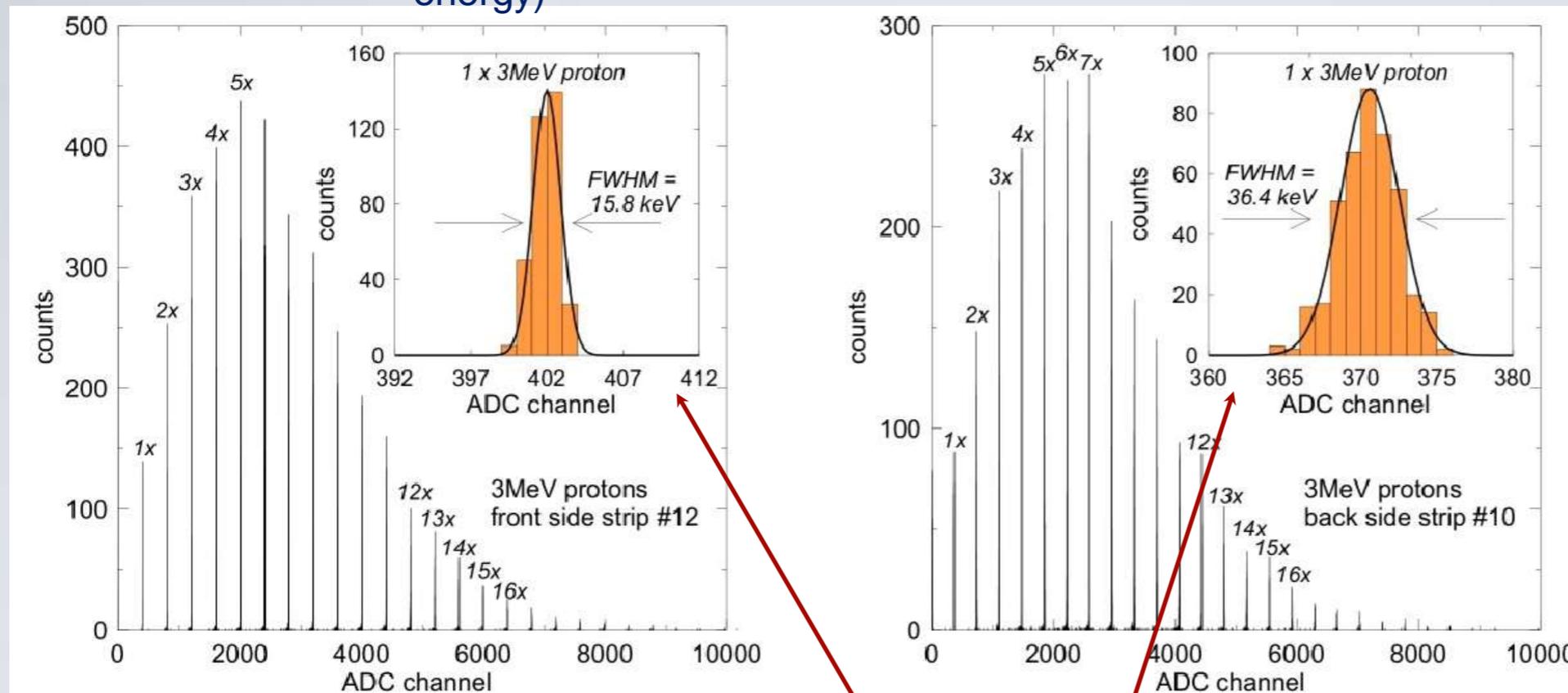
Si-CsI(Tl) matrix



Preliminary results... analysis underway
Ne + CH₂ → Multifragmentation

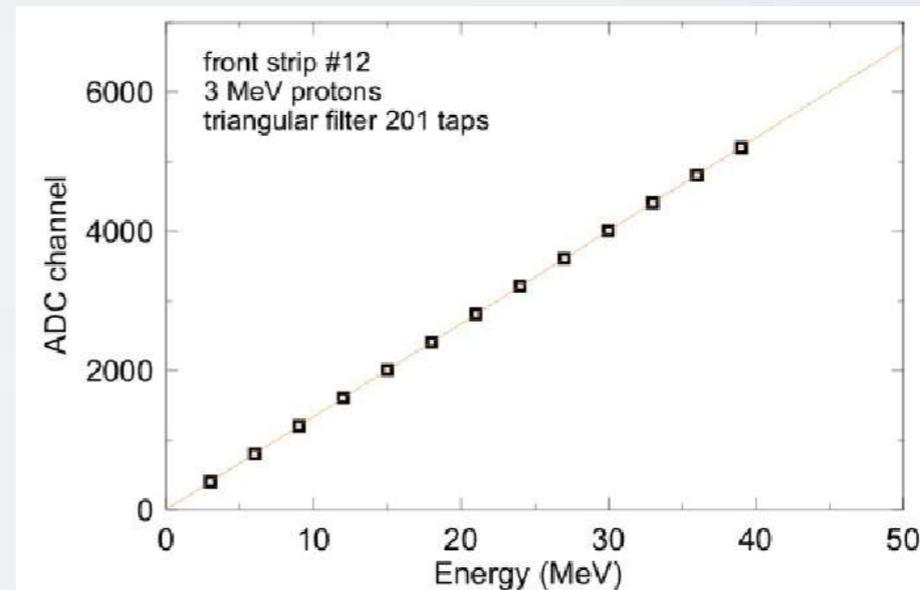
n-proton injection and resolution

Proton bunch multiplicity (deposited energy) →



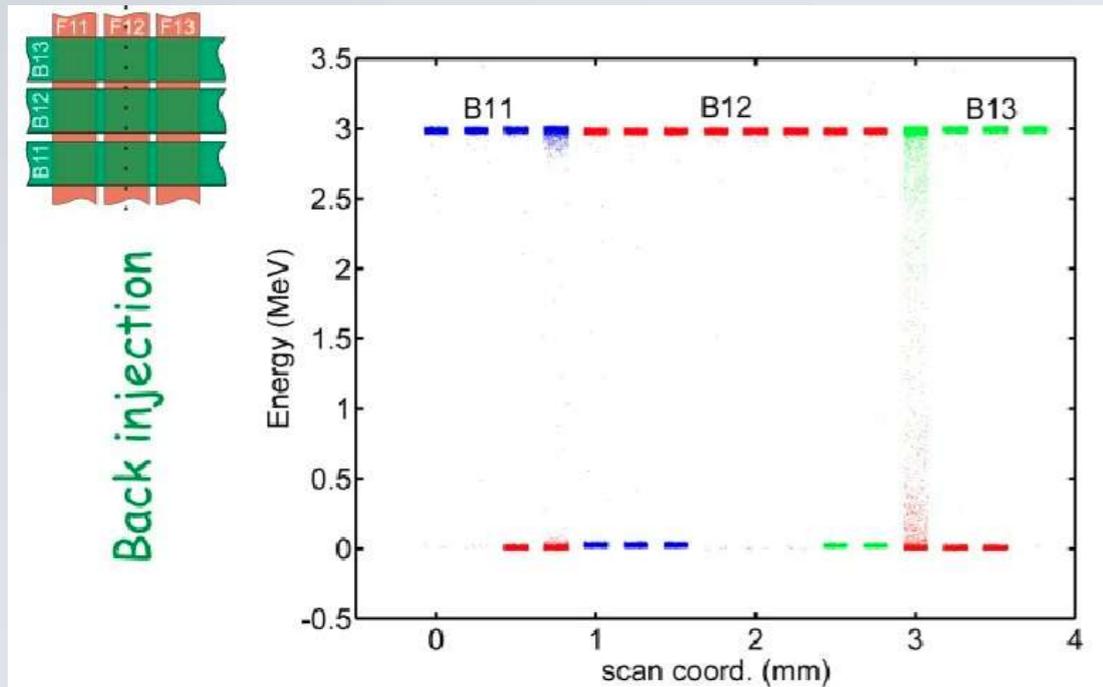
Resolution: 0.5% Front - 1.2% Back

Test of linearity for
Silicon+Preamplifier
system

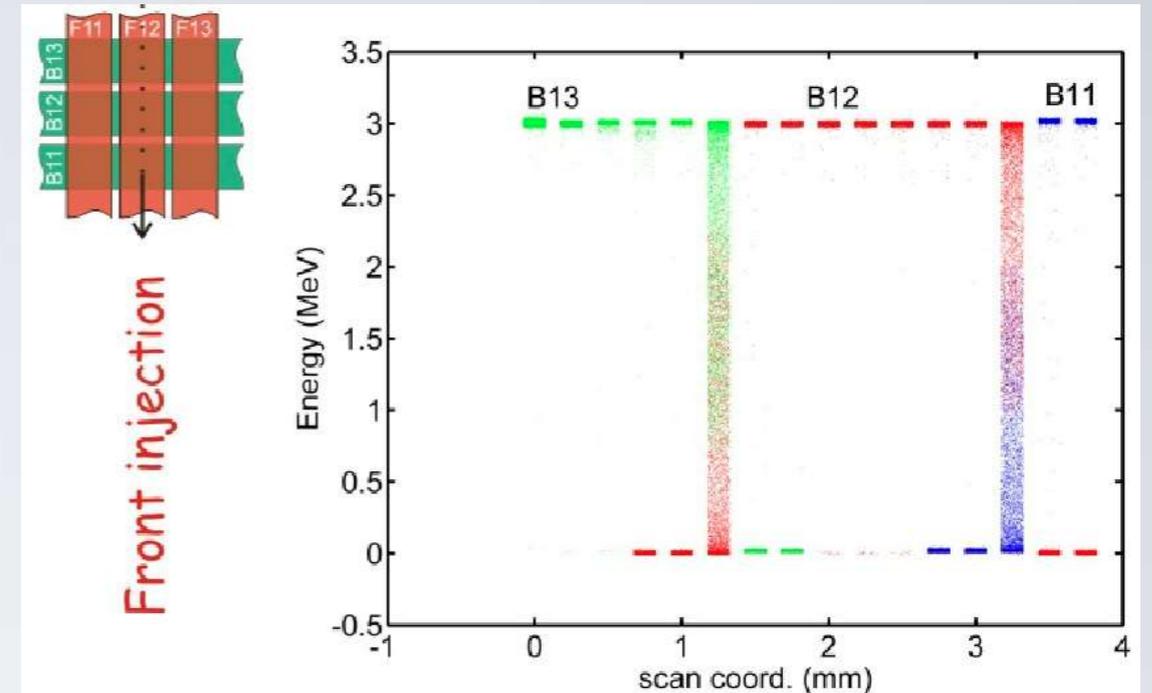


Interstrip energy sharing

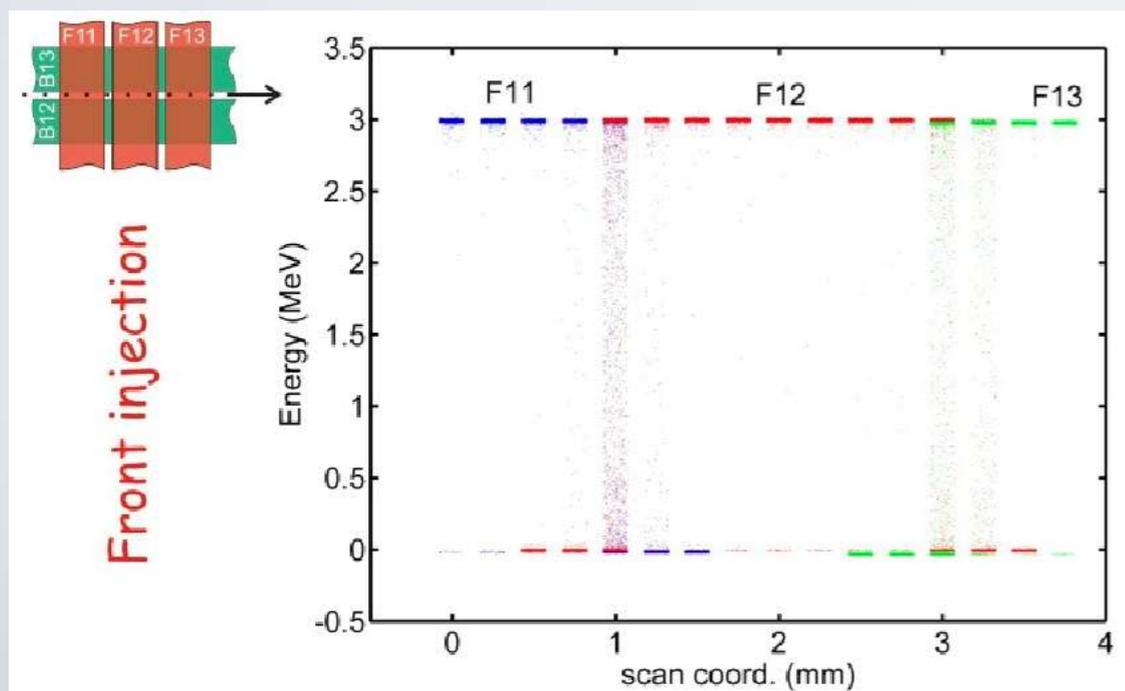
Back injection + Back strip/interstrip spanning



Front injection + Back strip/interstrip spanning



Front injection + Back interstrip spanning



Data analysis under way

*Study of energy sharing
between neighboring strips*

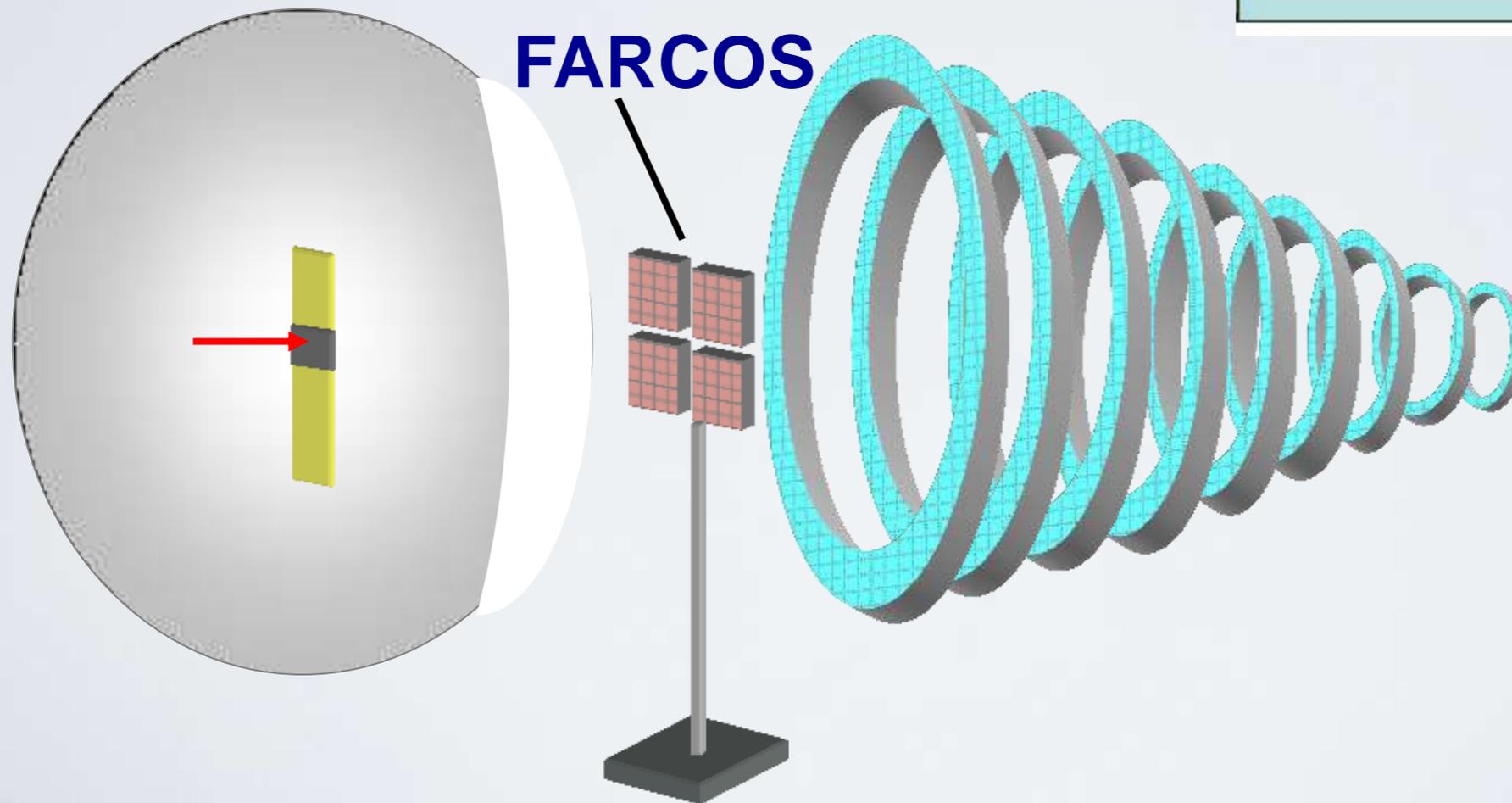
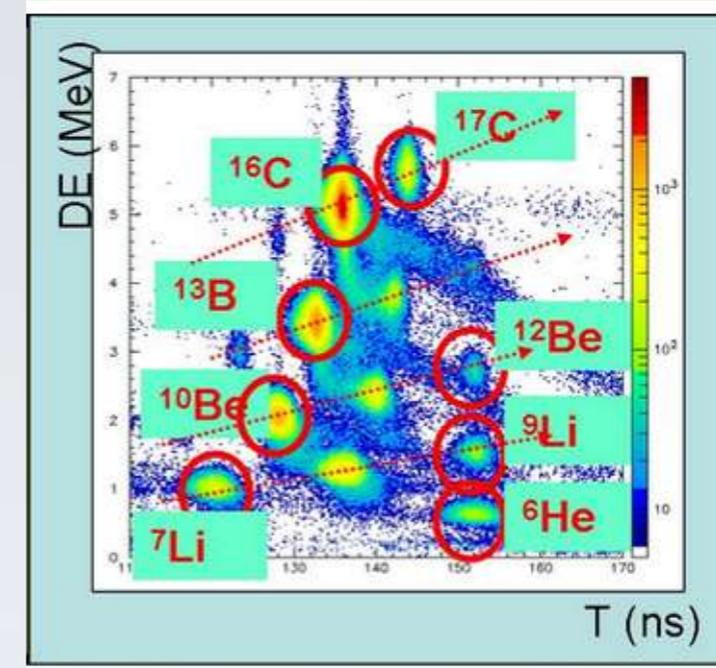
INFN and Politecnico,
Milano for Farcos
Collaboration

Experiments in 2013 @ LNS

CLIR: Clustering in Light Ion Reactions



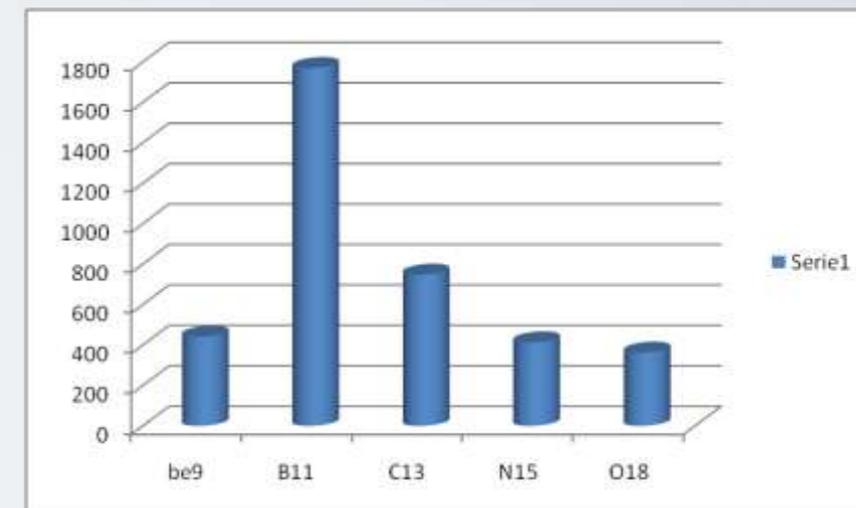
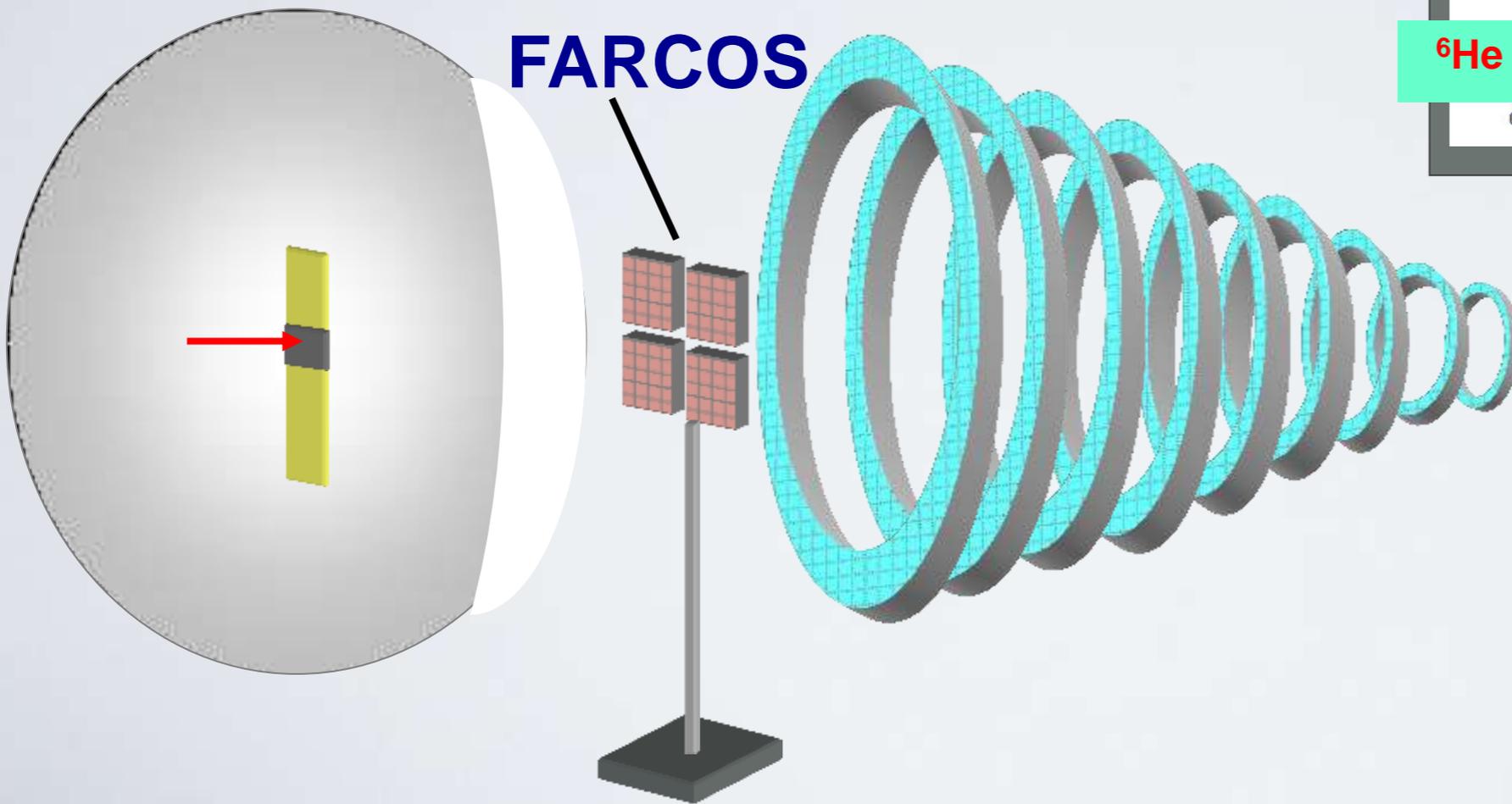
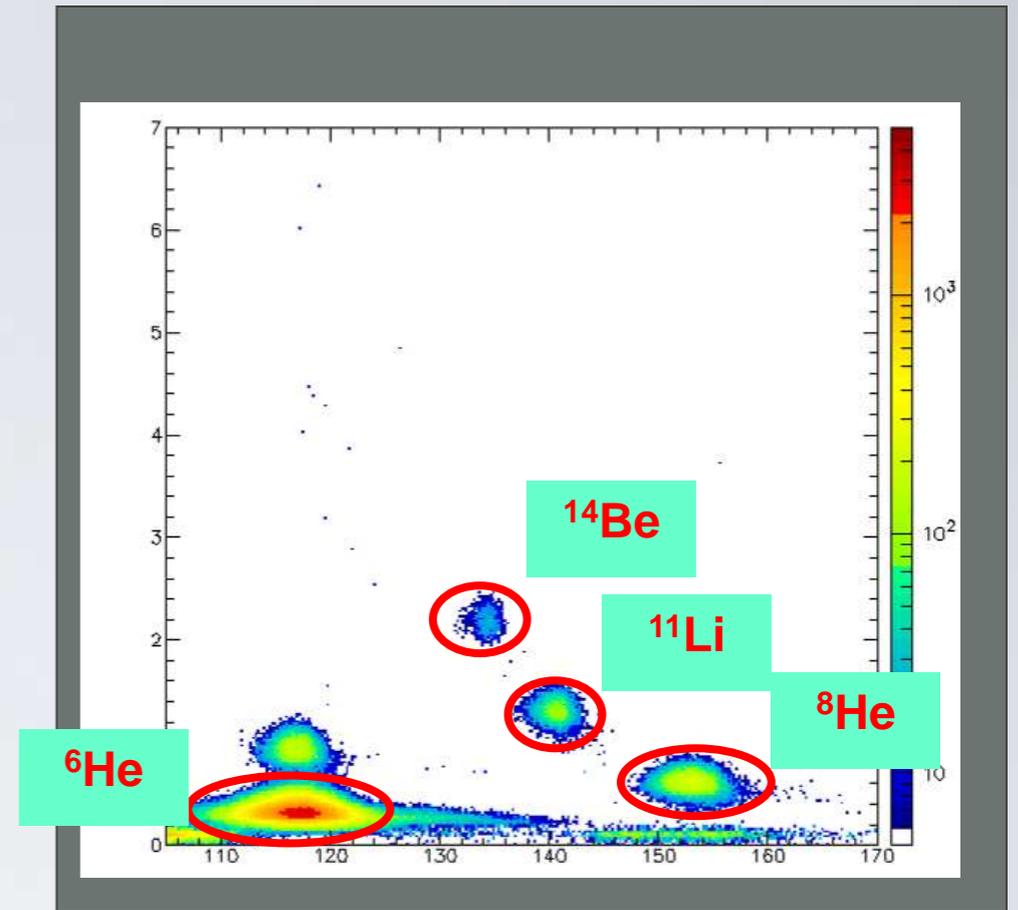
16C from tagged FRIBS projectile
fragmentation \rightarrow 120 kHz



Experiments in 2013 @ LNS

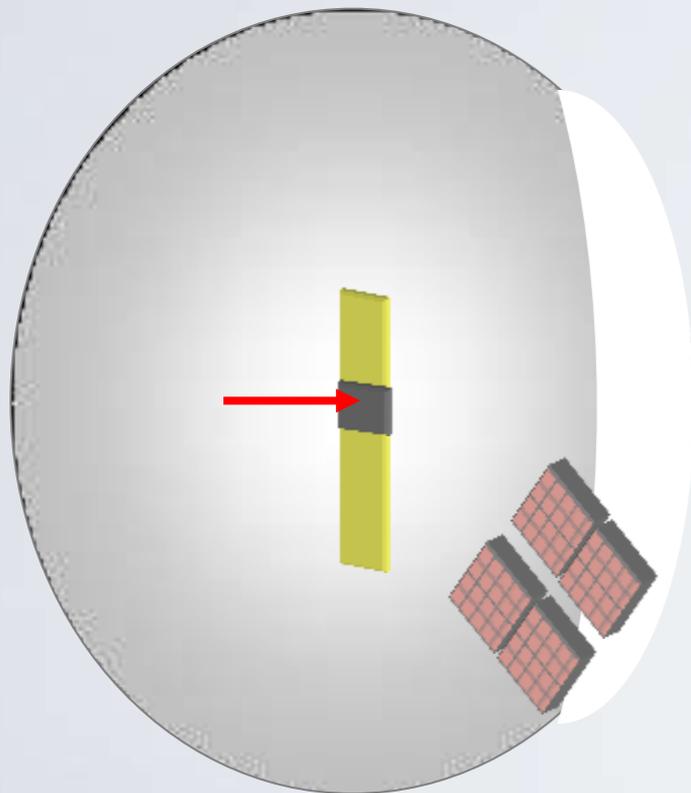
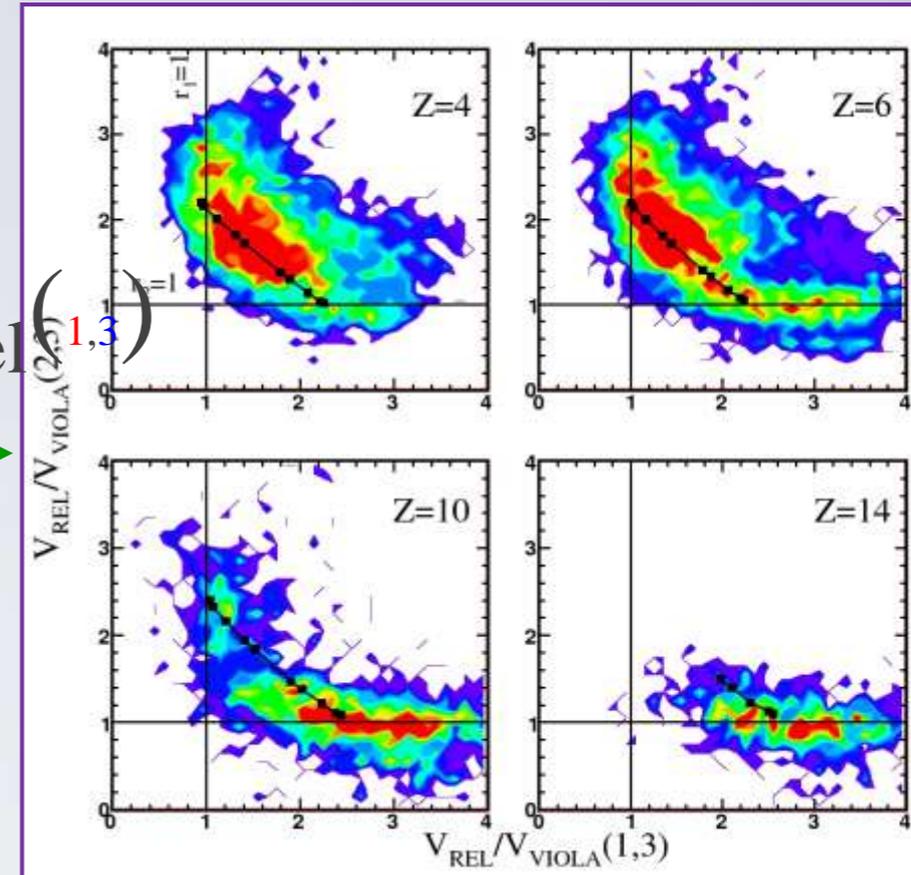
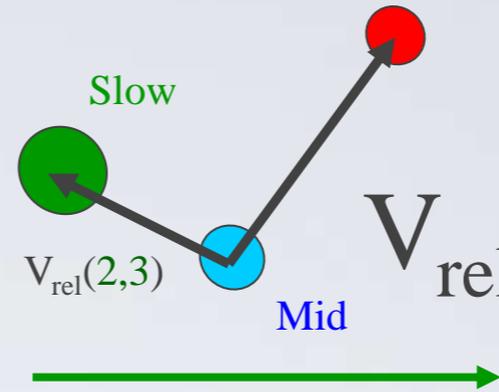
^8He : populations of resonances in ^9He
via d-p transfer reactions on ^8He

^{16}C from tagged FRIBS projectile
fragmentation \rightarrow 120 kHz

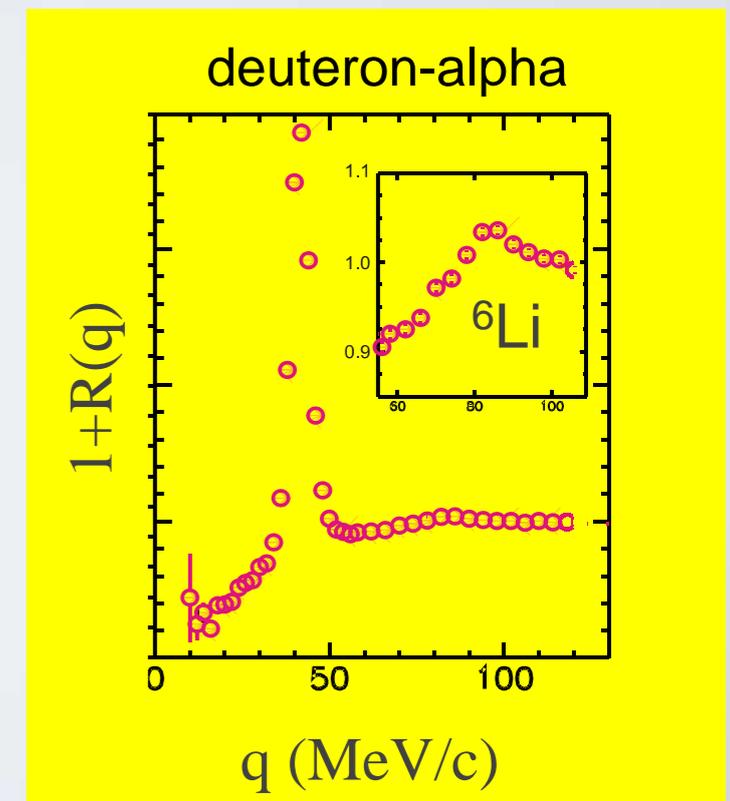
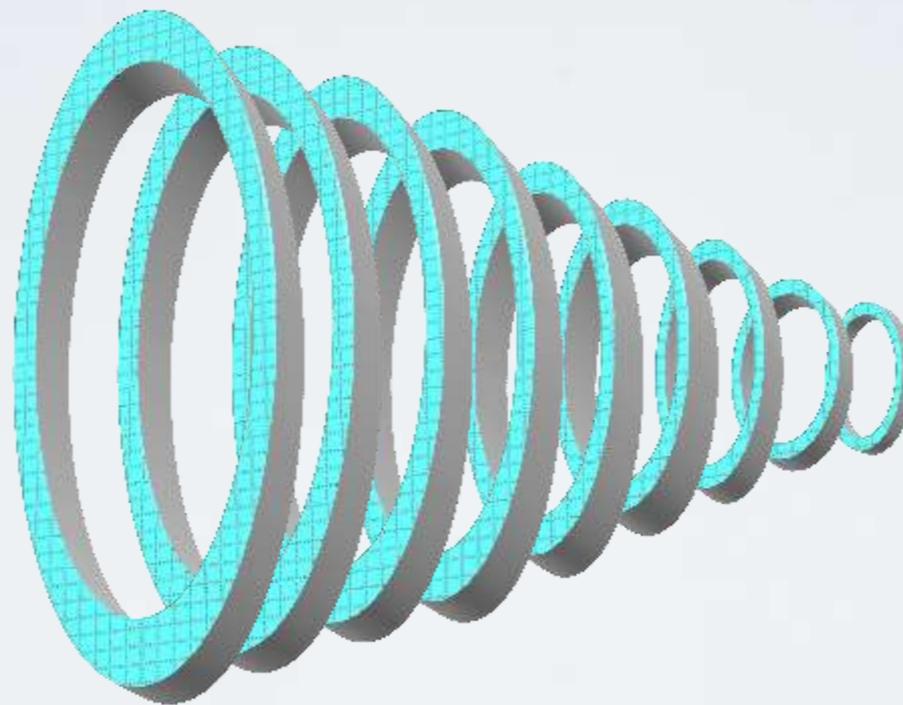


Experiments in 2013 @ LNS

inkinlsy mass and isospin dependence
 of reaction mechanism
 Comparison of old $^{124}\text{Sn}+^{64}\text{Ni}$ data with
 $^{124}\text{Xe}+^{64}\text{Zn}$
 Use of farcos for p-p correlation for the
 comparison of source size and life-time



FARCOS



Electronics hopes for phase 2

20 telescopes??

- Pre-amplifiers with multiple gains and large dynamic range (MeV to GeV)
- VSLI, configurability, low power... oriented towards physics!
- Pulse-shape capabilities
 - ✓ Low identification thresholds for low energy experiments
- Digitalization of detector signals
- Update possibilities
 - ✓ Solid angle increase (20 telescopes?)
 - ✓ Coupling to different detectors (in different laboratories)

Electronics hopes for phase 2 with GET

- 20 telescopes means – 2560 strips + 80 CsI channels
- Means (including spares) about :
- 60 AGET – 15 ASAD – 5 COBO – 2 MUTANT
- But preliminary test:
- Is the dynamical range and bit resolution enough?
- Do we need to use two cannels for each strip at least for front strips?
- Can you produce for us the electronics and what will be the price?
- Can we get the money from INFN?
- Can we couple the GET ACQ to CHIMERA ACQ?
- Fase 3?? Can we use GET or “upgraded-GET” to replace the old CHIMERA electronics (about 2400 channels CsI and silicon)?

Man power

- We are not too much
- For electronics:
- 1 staff engineer in Catania 3 students (Phd and postdoc levels) and some technicians no big experience in VLSI.
- Collaboration with 2 staff engineer and some students from Politecnico Milano with experience in VLSI.
- We are beginning some collaboration with Engineer department in Catania – experience on VLSI but not for nuclear physics.
- Some physicist experts in ACQ (1+1) and electronics (3 but 2 are the same of ACQ) help from some technicians.
- Available fellowships at INFN for engineers – possible collaborations with LNL expertise for SPES detector development.

Status and... coming up...!

- Farcos has a preliminary configuration allowing for experiments at LNS and other laboratories (ex. GANIL)
 - ✓ Compact Preamplifiers, shapers, etc. for about 800 channels
 - ✓ Limited digitalization capabilities already available
- The array is going to increase in size... need for integrated technology
 - GET solution after the front-end; new integrated front-end (GET, INFN Milan-Politecnico, ...) we have to see if possible before the end of this year