Transfer reactions on light exotic nuclei with CHIMERA detector at LNS

G.Cardella for the EXOCHIM collaboration



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Outline

Short presentation of the CHIMERA detector and LNS fragmentation beams

Can one get precise angular distributions without extremely segmented detectors?

≻Test case ¹⁰Be+p→⁹Be+d

> Extension of the method using coincidence γ -ray

Conclusions and perspectives

The CHIMERA detector : particle identification methods



Fragmentation beams at INFN-LNS - Catania



-Neutron transfer reactions near halo nuclei -

We study direct reactions using light exotic nuclei impinging on p, d targets useful to investigate on various structure effects (see Alexandre Obertelli this morning)

EVENT SELECTION performed with kinematic coincidences – measuring in binary/ternary reactions all reaction partners we clean the events



– Advantages of binary kinematics : the ¹⁰Be+p→⁹Be+d case -



Kinematical coincidence method in transfer reactions

L. Acosta^b, F. Amorini^b, L. Auditore^d, I. Berceanu^h, G. Cardella^{a,*}, M.B. Chatterjieeⁱ, E. De Filippo^a, L. Francalanza^{b,c}, R. Giani^{b,c}, L. Grassi^{a,k}, A. Grzeszczuk^j, E. La Guidara^{a,g}, G. Lanzalone^{b,e}, I. Lombardo^{b,f}, D. Loria^d, T. Minniti^d, E.V. Pagano^{b,c}, M. Papa^a, S. Pirrone^a, G. Politi^{a,c}, A. Pop^h, F. Porto^{b,c}, F. Rizzo^{b,c}, E. Rosato^f, P. Russotto^{b,c}, S. Santoro^d, A. Trifirò^d, M. Trimarchi^d, G. Verde^a, M. Vigilante^f



The lab energy of the detected particle determines the CM emission angle

Due to the relatively good energy resolution we can obtain an angular distribution with much better resolution than the one determined by the size of the detectors

– STEPS of the analysis ¹⁰Be+p→⁹Be+d

We select only complete events with two detected particles and with total detected charge Ztot=Zbeam+1

We can plot the $\Delta \phi$ angle between the two coincidence detectors – due to momentum conservation $\Delta \phi$ must be 180°

 $\Delta \phi$ width due to the finite opening of the detectors





We also clean the events putting constraints on the total detected energy must be equal to the beam energy 580 MeV + Q_{value} (-4.58 MeV)

Notwithstanding the scarce total energy resolution we see only GS events ${}^9\text{Be}^* \rightarrow n+\alpha+\alpha$ S_n=1.66 MeV



Fig. 5. Relative angle $\Delta \phi$ between the telescopes selected in coincidence in the reaction ${}^{10}\text{Be}+\text{p} \rightarrow {}^{9}\text{Be}+\text{d}$. The peak at 180° is due to kinematical coincidences.

- The ¹⁰Be+p \rightarrow ⁹Be_{q.s.}+ d angular distribution -



Note that angular distributions are automatically corrected for the fragmentation beam angular spread

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- Excited levels - γ-ray tagging? -

counts



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– Using the γ -ray tagging: the ¹⁰Be+p \rightarrow ¹⁰Be+p case -



The angular distribution we measure for this channel shows a slope change around 40°

This change of slope is not seen in previous data measured by Cortina-Gil et al Phys.Lett.b 401(1997)9

The proton angular distribution measured in coincidence with γ -rays explains the change of slope and the deviation from Cortina-Gil data

	E _{level} (keV)	Jп	T _{1/2}	E _v (keV)	I _Y	γ mult.	Final leve	1
	0.0	0+	1.51E+6 y 4 % β ⁻ = 100					
	3368.03 <i>3</i>	2+	125 fs 12	3367.415 30	100	E2	0.0	0+
	5958.39 5	2+	<pre>% II = 100 < 55 fs % IT = 100</pre>	2589.999 60 5955.9 5	>90 <10	M1 E2	3368.03	2+ 0+
	5959.9 <i>6</i>	1-	% IT = 100	2591.5 6 5958.0 6	17 8 83 8	E1 E1	3368.03 0.0	2+ 0+
	6179.3 7	0+	0.8 ps +3-2 % IT ≈ 100	219.4 3 2811 7 6178	24 2 76 2	E1 E2 E0	5959.9 3368.03 0.0	1- 2+ 0+
	6263.3 50	2-	% IT = 100	303.4 50 2894.9 50 6261.2 50	≤1 99 1 1 1	E1 M2	5959.9 3368.03 0.0	1- 2+ 0+

The γ-ray spectrum measured is what we expect from ¹⁰Be* excited states decay

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Future improvement : FARCOS as forward angle spectrometer

New triple telescopes Si-Si-Csl First stage 32x32 strip 6.2x6.2cm² 300 μm DSSSD Second stage 32x32 strip 1500 6.2x6.2cm² μm DSSSD Third stage 4 Csl(tl) 3.1x3.1 cm² photodiode readout 6 cm thick First test-experiment coupled with Chimera April 2013 see E.De Flippo next Thursday C6 and Acosta poster Nf001



Conclusions and perspectives

Using cocktail of neutron rich beams with the CHIMERA detector we are able to extract angular distributions for many reaction channels searching for structure effects on cross sections

The 4π detection efficiency is very useful and allows extensive use of the kinematical coincidence technique

We can also detect γ -rays with our CsI(TI) detectors in order to tag excited levels

For the future experiments we are working to improve our detection capabilities and resolutions also coupling CHIMERA to a new high resolution strip telescope array FARCOS and to neutron detectors



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L.Acosta¹, F.Amorini^{1,3}, A.Anzalone¹, L.Auditore⁴, I.Berceanu⁹, G.C.², M.B.Chatterjee¹⁰, E.De Filippo², L.Francalanza^{1,3}, S.Gianì^{1,3}, L.Grassi^{2,3}, E.La Guidara^{2,5}, G.Lanzalone^{1,6}, I.Lombardo⁸, D.Loria⁴, E.Morgana⁴, T.Minniti⁴, A.Pagano², E.V.Pagano^{1,3}, M.Papa², S.Pirrone², G.Politi^{2,3}, A.Pop⁹, F.Porto^{1,3}, F.Rizzo^{1,3}, E.Rosato⁸, P.Russotto^{2,3}, S.Santoro⁴, A.Trifirò⁴, M.Trimarchì⁴, G.Verde², M.Vigilante⁸ 1) INFN Laboratori Nazionali del Sud, Catania, Italy 2) INFN, Sezione di Catania, Catania, Italy 3) Dipartimento di Fisica e Astronomia Università di Catania, Italy 4) INFN gruppo collegato di Messina & Dipartimento di Fisica e Astronomia Università di Messina, Italy 5) Centro Siciliano di Fisica Nucleare e Struttura della Materia, Italy 6) Università Kore di Enna, Enna, Italy 7) INFN Sezione di Milano & Dipartimento di Fisica e Astronomia Università di Milano, Italy 8) INFN Sezione di Napoli & Dipartimento di Fisica e Astronomia Università di Napoli, Italy 9) Institute for Physics and Nuclear Engineering, Bucharest, Romania

10) Saha Institute of Nuclear Physics, Kolkata, India